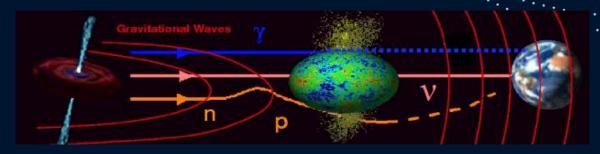


Status and physics results of the KM3NeT experiment

A. Chen
University of the Witwatersrand
On behalf of the KM3NeT Collaboration



NEUTRINOS AS COSMIC MESSENGERS



Neutrinos are neutral, stable, weakly interacting

- not absorbed by background light/CMB
- not absorbed by matter
- not deviated by magnetic fields

access to dense environments

access to cosmological distances

astronomy over full energy range

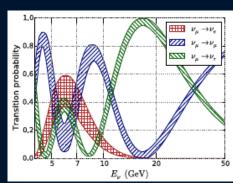
'Smoking gun' signature for hadronic processes

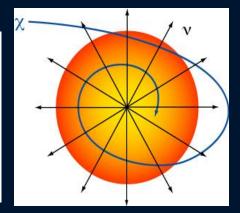
Correlated in time/direction with electromagnetic and gravitational waves

NEUTRINO ENERGY RANGE

MeV to PeV energies









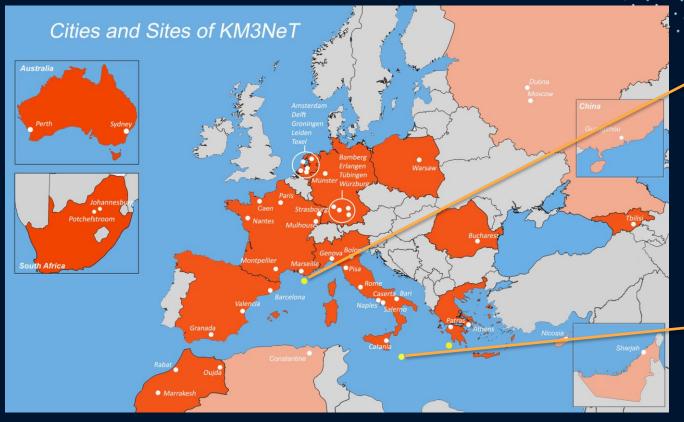
Supernovae Solar flares

Atmospheric v v oscillations v mass ordering sterile v

Dark matter Monopoles, Nuclearites,...

Cosmic v Production mechanism of HE CR

THE KM3NeT EXPERIMENT





Oscillation Research with Cosmics In the Abyss

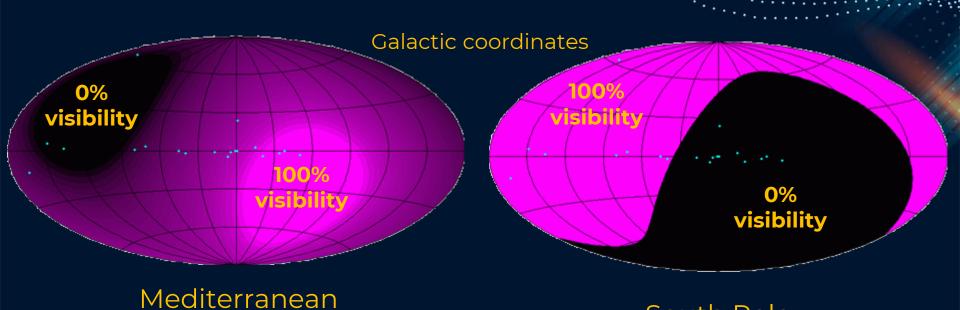


Astroparticle Research with Cosmics In the Abyss

4

WHY IN THE MEDITERRANEAN SEA?

~ 43° North



South Pole

THE KM3NeT DETECTOR

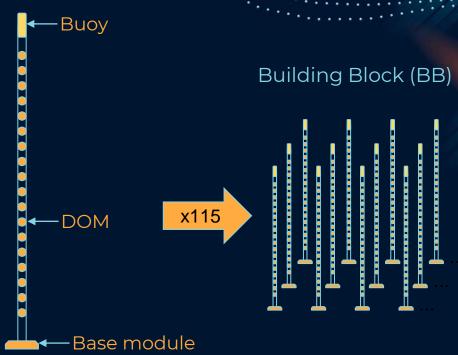
Digital Optical Module (DOM)



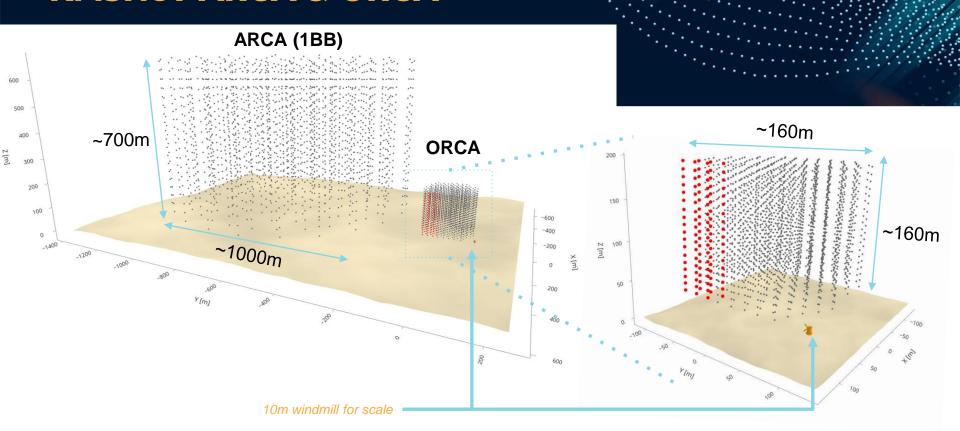
x18

- 31x3" PMTs
- ns timing
- ~10 cm spatial positioning

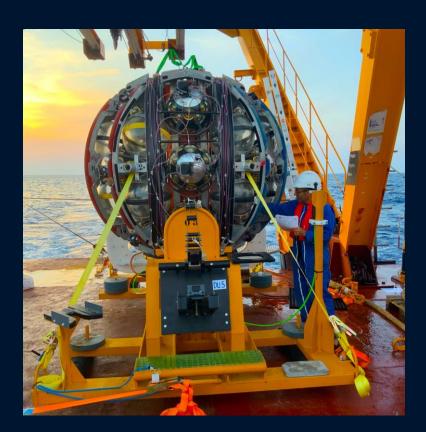
Detection Unit (DU)

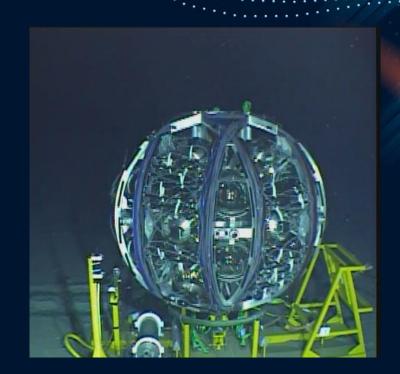


KM3NeT ARCA & ORCA



THE DETECTOR CONSTRUCTION





THE DETECTOR STATUS

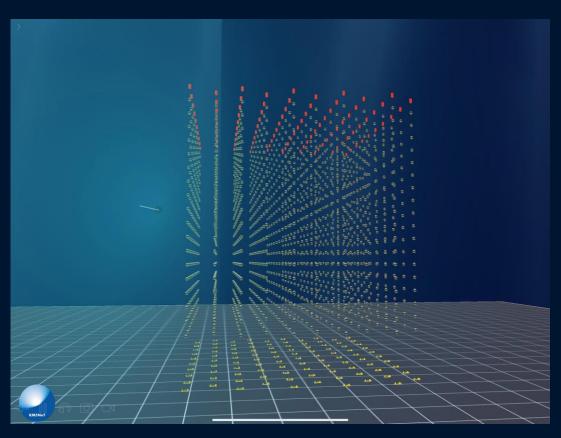
KM3NeT-ARCA

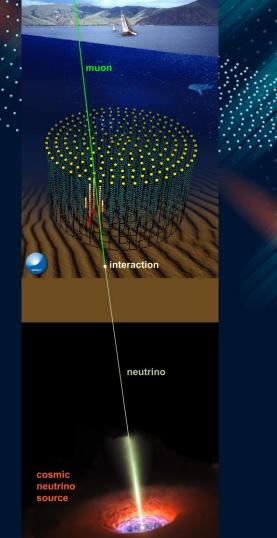
8 DU currently deployed

KM3NeT-ORCA

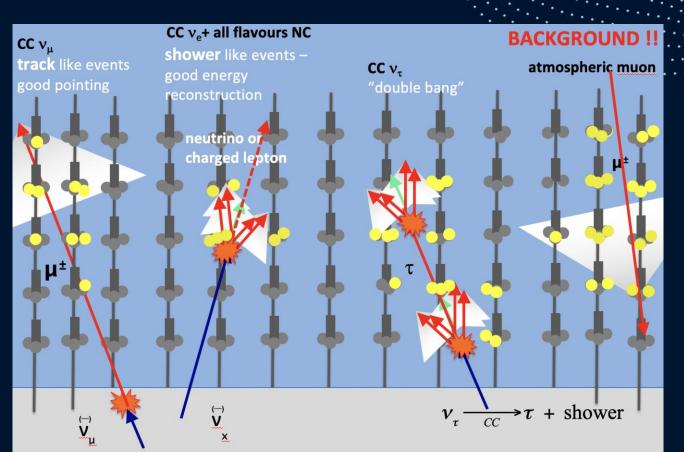
10 DU currently deployed

THE DETECTION PRINCIPLE

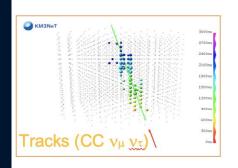




EVENT TOPOLOGIES

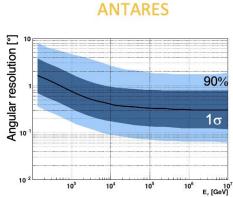


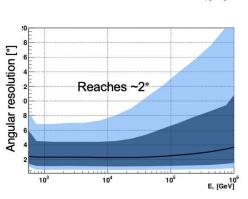
RESOLUTIONS

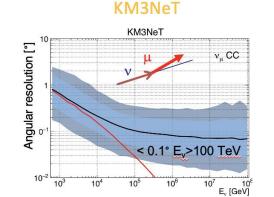


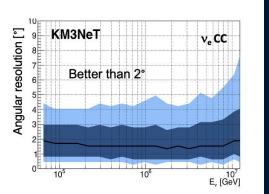
Cascades (CC ve vt - NC)

◯ KM3NeT







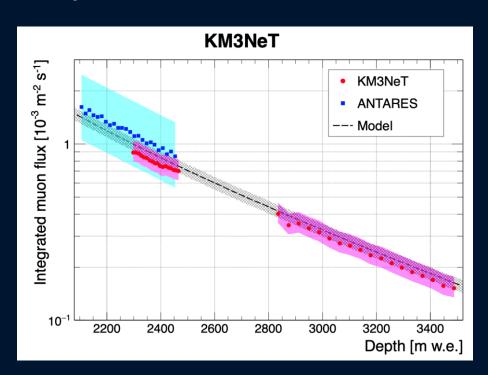


Energy resolution ~ 30%

Energy resolution < 5%

FIRST RESULTS

Atmospheric muons



ARCA 2 (23/12/2016-2/3/2017) ORCA 1 (9/11/2017-13/12/2017)

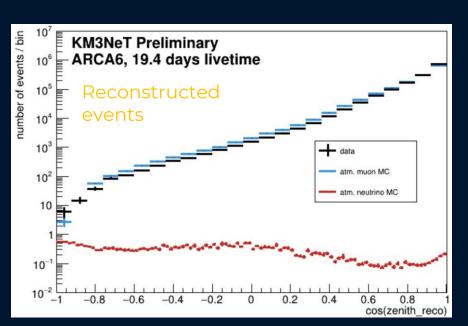
Muon flux as function of depth compared to Bugaev model

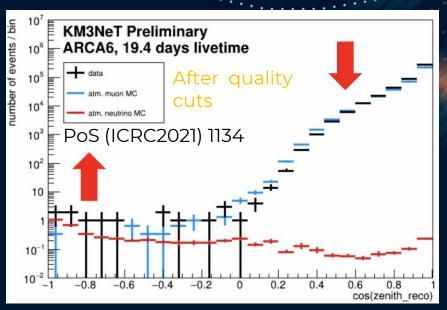
(Bugaev et al, Phys. Rev. D 58 1998 054001)

Eur. Phys. J. C 80 (2020) 99

FIRST RESULTS

Atmospheric neutrino





Upgoing tracks = neutrino induced events Downgoing tracks = dominated by atm muons 15 up-going tracks > 100 GeV MC expectations: 4 atm ν + 7 atm μ

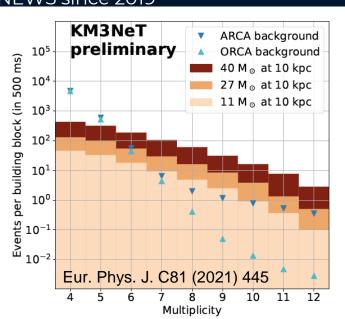
SUPERNOVA MONITORING

Neutrinos below 100 MeV expected at several stages of the core collapse

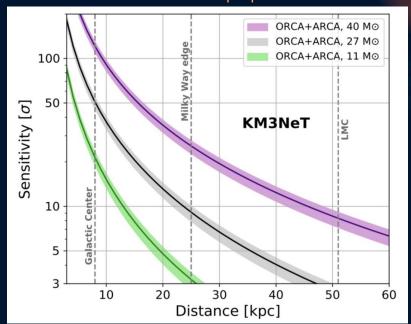
Cherenkov signature detected as a population

of coincidences in single DOMs

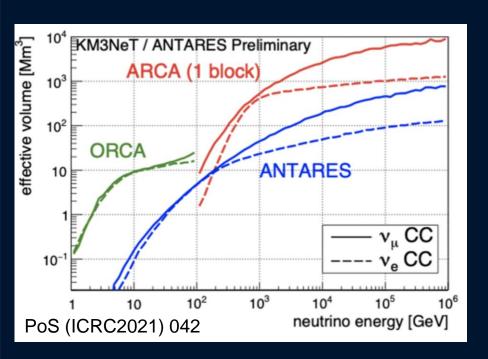
Real time alert system is in place within SNEWS since 2019

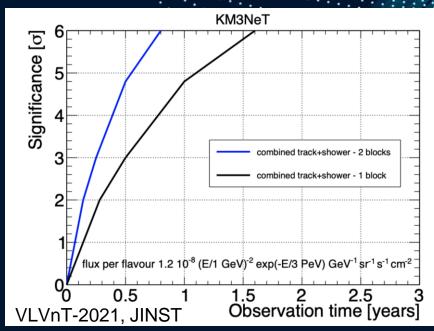


 5σ for ARCA+ORCA for $27M_{\odot}$ at a distance ~ 36 kpc ~60% of the CCSN Galactic population



DIFFUSE COSMIC NEUTRINO

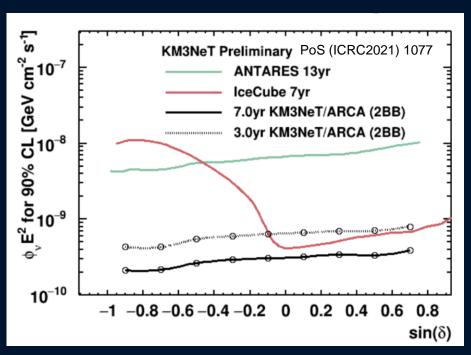


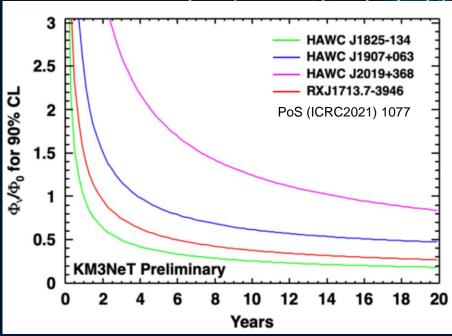


Broad energy coverage of the full detector

 5σ detection in less than 1 year for 1 BB!

POINT-LIKE NEUTRINO SOURCES

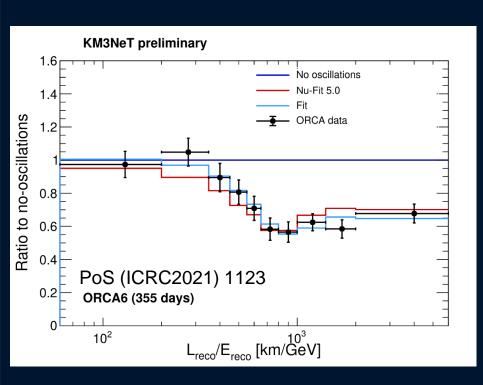


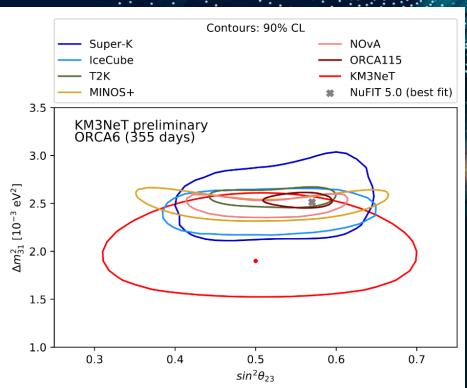


E⁻² neutrino flux 90% CL sensitivity

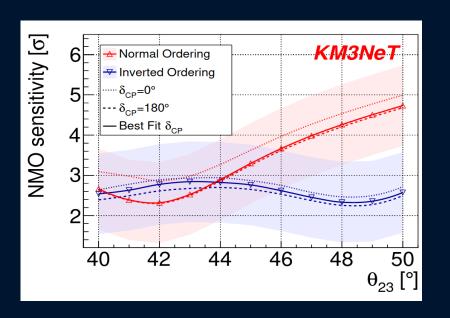
90% CL sensitivity for several candidate sources

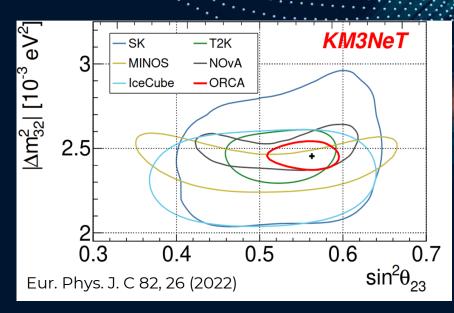
NEUTRINO OSCILLATION





NEUTRINO MASS ORDERING

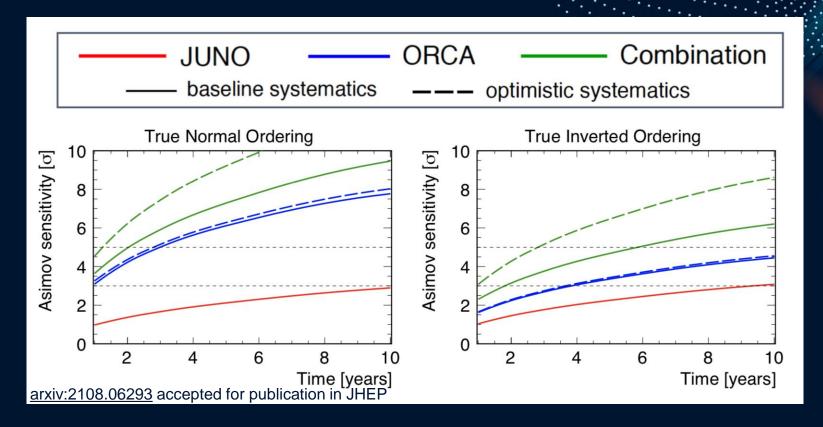




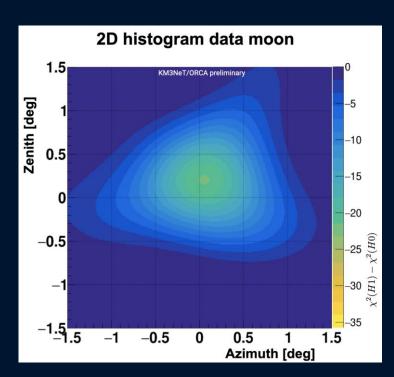
Expected results for 3 years exposure, full detector.

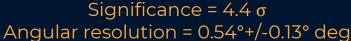
Competitive sensitivity to Δm^2_{32} , θ_{23}

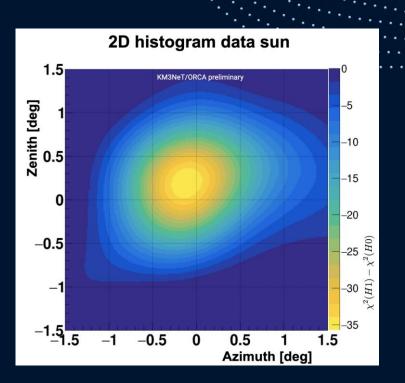
NMO COMBINED WITH JUNO



MOON/SUN SHADOW SENSITIVITY







KM3NeT ORCA 13 months data taking

Significance = 5.7σ Angular resolution = 0.59° +/- 0.10° deg

CONCLUSIONS

- KM3NeT Neutrino telescopes will allow neutrino physics and neutrino astronomy from the MeV scale to the PeV scale
- Well established underwater technology and calibration procedures
- First preliminary results confirm expectations
- KM3NeT effective area already larger than ANTARES
- 8 ARCA + 10 ORCA Detection Units have been deployed
- More Detection Units to be deployed this year

STAY TUNED!



THANKS FOR YOUR ATTENTION