

### Micro Pattern Gas Detectors

- CERN workshop introduction
- Gaseous detector introduction
- First MPGD
- First evolution  $\rightarrow$  more robust devices
- Second evolution  $\rightarrow$  Introduction of resistive layers
- MPGD selection guide
- Future tasks



- -21 persons
- -14 persons from a contracting company
- -7 CERN staff
- -1400 square meters (new building operational in 2018)
- -Making PCBs since 1965
- -Mandate:
  - -produce parts difficult to find in industry
  - -provide to CERN users expertise in the field
  - -make some mass productions



Etc..



# Introduction to Micro Pattern Gas Detectors

### Gaseous detectors principle



## Gaseous detectors principle



### Gaseous detectors principle



# Till 1988 Wire chambers







#### Wire chambers

- Slow ion evacuation.
- Pitch limited to 1 to 2mm causing limited R/O granularity.
- Classical mechanical production techniques.
- Exponential gain, most of the amplification is done in the last micro meters.
  Drift and amplification regions are defined by the field shape.
- Possibility of aging and polymerization.



#### • 1988 First MPGD

# Micro Strip Gas Chamber (MSGC)

#### MSGC

Anode strips of 10 um with a pitch of 200 µm on a glass substrate.

Spatial resolution ~50µm Rate capability ~10<sup>6</sup> Hz/mm2

- Faster ion evacuation
- Higher spatial resolution
- 100um scale structures
- Photolithographic techniques
- Still exponential gain close to the thin anode
- Drift and amplification field are separated





#### But this device shown some problems in HEP environment



-less than 1um thick metal

-Mandatory thickness to keep the sub-micrometer precision at the Strips edges

#### Spark damages



-Triggered by heavy ionizing particles.

-Unstoppable erosion of electrodes.

-Spark energy above the materials capabilities to evacuate it .

-This triggers : evaporation, metal melting, carbonization. Sign of really high temperature processes

#### Aging/polymerization



-Creation of surface or dendritic deposits due to complex gas molecules in presence of exponential fields.



- Nineties  $\rightarrow$  First evolution
- Explosion of ideas to solve these 2 defects



#### 

#### It took a decade, but at the end, 3 structures have emerged!



The problem of aging and polymerization have been solved by removing exponential fields, in addition to a precise materials and gas selection. Such problems have not been seen anymore.

The detectors were more robust against sparks thanks to thicker conductors, but not completely immune.

It is interesting to observe that in these 3 options the amplification gaps are precisely defined by industrial materials, not by photolithography: -Polyimide 50um +/-0.1um -Photo-imageable coverlay spacer 128um +/-2um

## 1997→Micromegas (Ioannis Giomataris)



- -Electroformed meshes, mechanically fragile.
- -Spacers made with fishing wires.
- -Delicate artisanal production.
- -Special anode strip connection to FE to minimize sparks damages.
- -The use of this structure stayed quite confidential till 2007 due to production difficulties.

#### 1998 $\rightarrow$ GEM foil and then triple GEM detector (Fabio Sauli)







Spark problem solved: →triple GEM structure

Aging problem solved: →no exponential A-field →Specific gases →Specific materials

The amplification gaps are not affected by the detector size.

A GEM is only made with Photolithographic processes. Industrial by nature.

The technology grow was fast.

#### GEM

•Base material : Polyimide 50um + 5um on both sides



Limited to 40cm × 40cm due to:
The 2 masks alignment precision
And Glass mask cost



Limited to 2m × 60cm due to:
Base material
Equipment

#### Double mask triple GEM detector















COMPASS Settled all the parameters: -GEM Segmentation -HV distribution -X/Y Read out board -FE electronics



Spherical GEM R&D for X-Ray diffraction application

# "Single mask" was introduced to make larger detectors and reduce the cost



1.8m





A 10cm × 10cm double mask GEM costs 300 \$/piece.

For 100 single mask GEM the cost drops to 100 \$/piece.

For 10 000 pieces , less than 50 \$.

There is ways to even go below for Larger quantities!

# CMS upgrade with single Mask



CMS nose

### Other detectors with single mask technique



Future CMS MEO  $\rightarrow$  1000 GEMs

KLOE - Cylindrical Detector

## And many more

-BM@N in Dubna (1.6m x 0.5m) -SBS tracker Jefferson lab -CBM at Fair -BESIII China

-SOLID -BONUS 12 -P-RAD -S-Phenix TPC -COMPASS upgrade

-GEM for nuclear physics TPCs

-ESS for neutron detectors

-and lot of small GEMs for academic purpose

ALICE TPC  $\rightarrow$  700 GEM

#### 1998→first LEM or THGEM (Pio Picchi)

Mechanical Drilling : FR4, PMMA etc.. More recently: Sand blasting on glass Photo imaged glass







Eltos (IT) Hold by Fulvio Tessarotto



Print electronics (IS) Hold by Amos Breskin



CERN Hold by Serge Ferry

#### **RIM** creation



#### Other types of THGEMs









# 1999→Micro-Well (Ronaldo Bellazzini)



-3 x 3 cm Micro-well detector -Produced at MPT with GEM processes -Really simple but abandoned due to the difficulty to mitigate sparks



Close-up view Square pattern used in the early days







#### $2007 \rightarrow A$ better process was found to produce MM



BULK Micromegas

473.0



#### Floating mesh Micromegas



The detector behavior was really good ! The production was easier, repeatable and using industrial processes. The technology spread was then rapid.



#### Read-out board

400um pillar





Pillar matrix on 1cm x 1cm pads





Mesh stretching & gluing on frame



#### Coverlay lamination



Detector developed / cured / tested

## **BULK Micromegas detectors**



T2K TPC ,J.Beucher 1.8m × 0.8m plane With 12 detectors



ILC DHCAL , M.Chefdeville 1m x 1m plane With 6 detectors







Early ATLAS NSW R&D Joerg Wotschack 1.5m x 0.5m plane Single panel





33 sectors , 12cm diametertest detector2.5mm dead space for sectorization1mm hole for HV connection

- 2012 Second evolution
- Resistive layer introduction.
   A long path, I think not yet fully explored !



In 2012 during the early phase of ATLAS NSW R&D we faced a new problem.

With LHC background the single amplification stage large MPGD's were continuously sparking, they were not damaged but simply constantly stopped due to the spark rate

In close collaboration with <u>Joerg Wotschack</u> we decided to introduce resistive layers.

It took us 12 iterations to understand how to solve the problem adding resistive layers

#### First but bad idea : Cover the R/O lines with Resistive material



# And finally, after 4 more trials!



#### Validation with Neutron irradiation

Without resistive layer

#### With resistive



#### ATLAS NSW prototype $1m \times 2m \rightarrow Floating mesh$



Open





#### Closed







### Atlas NSW



Joerg Wotschack



Close to 2000 Micromegas detectors produced with modules sizes up to  $2m \times 0.5m$ 

PCBs with pillars built at ELTOS (IT) and ELVIA (FR) Panels construction and detector Assy :

-Dubna -INFN Frascati -CEA Saclay -LMU Munich MPT participated to the R&D and was also involved in the mass production with industry -Specification -Companies selection -Technology transfer

### Right after we did a Survey on existing resistive layers to take a direction for the future

Polymer paste 10Kohms to 100K/Square

RuO Thick film paste 10K to 100M/square

Resistive ceramics or glass

Resistive Kapton

Dissipative films

DLC: Diamond Like Carbon

Too low values (but strips can artificially increase the value)

Perfect range , but a ceramic substrate is mandatory

Limitation on resistive values and size

Limitation on resistive values

Too high values

Perfect range 100K to  $1T \rightarrow$  chosen direction

#### A Large collaboration inside RD51 was raised :



Study: DLC deposition Detectors production Detectors characterization Aging

# DLC producers





Kobe





Lanzhou institute



And now CERN

Axis

700-800

600-700

500-600

400-500

300-400

200-300





Initial results: example of 1m x 0.6m foils 500Kohms/square +/-60% The error dropped today to values around +/- 30%

#### First benefit coming with resistive layers : Spark protection

- Spark energy :  $\frac{1}{2}$  CV2
  - 600 uJ in a 10cm × 10cm GEM.
  - 30 uJ with a classical MM 10cm x10cm.
  - 0.06 uJ (10M resistive layer).

#### First benefit coming with resistive layers : Spark protection





Massive structural effects Metal melting Kapton evaporation Processes happening above 1000 deg

But thanks to Triple GEM structure This is not happening

.40

#### First benefit coming with resistive layers : Spark protection

No structural defect No process above 100deg No effect on metals or Polymers

- Spark energy : <sup>1</sup>/<sub>2</sub> CV2
  - 600 uJ in a 10cm x 10cm GEM. 🗲
  - 30 uJ with a classical MM 10cm x10cm.
  - 0.06 uJ (10M resistive layer).
- Below 10M/square.
  - We can still create some visible compounds.
  - There is still enough energy to slightly deteriorate dielectrics.
- Above 10M/square
  - No visible compounds.
  - The structure is never damage.
  - Low Humidity and High cleanliness is mandatory !



Massive structural effects Metal melting Kapton evaporation Processes happening above 1000 deg

But thanks to Triple GEM structure This is not happening

• FE electronics input protections are not needed anymore

## second benefit : Resistive spreading

spice simulation Virginia university collaboration



Pulse amplitude and time shift with a 1MOhms/SQR layer

1600 cells 0.125mm x 0.125mm

# Resistive spreading



# Resistive spreading



# Third benefit : Possibility to do capacitive sharing

- -This is the solution to keep a large spreading keeping high resistive value !
- -The pulse is transmitted without delay.



#### uRwell with CS sharing layer 0.8mm pitch X/Y R/O



Performance of a resistive micro-well detector with capacitive-sharing strip anode readout Kondo Gnanvo<sup>a,\*</sup>, Nilanga Liyanage<sup>a</sup>, Bertrand Mehl<sup>b</sup>, Rui de Oliveira<sup>b</sup> <sup>1</sup> University of Virginia, Department Of Physics, Charlottesville VA 22903, USA <sup>6</sup> CERN Septemade des Particules 1P.0. Box 1211 Deneva 23, Switzerland

#### Fourth benefit : E-cleaning



Detectors with Kapton foils protecting the active area



Oven set at 90 deg



HV Power supply

1/Minimum requirement → the detector should hold 1/3 of final voltage in open air, leakage current <100nA.</li>
2/The active area should be covered with a Kapton foil to avoid dust dropping on.
3/The detector is put in the oven at 90 to 100deg Celsius.
4/A precise sequence is then applied to reach HV max (can be fully automatized).

Depending on the detector size and the type of defect the E-cleaning can takes few days.

#### Cleaning process



To avoid any re-deposition or nasty compound creation an extremely low level of humidity is required during this process.

### 2013→Resistive layer applied to MM structures



Damien Neyret Max Chefdeville Renaud Gaglionne Mauro Iodice Saclay Lapp Annecy Lapp Annecy INFN Roma



#### 2015→Resistive layer applied to Micro-Well (Giovanni Bencivenni) → uRwell

After vacuum gluing





10cm × 10cm µRwell detector "STD kit"

#### High rate uRwell

#### Charge evacuation in the active area







#### Best compromise between performances & cost $\rightarrow$ PEP uRwell



#### $P \rightarrow$ Pattern top copper

#### $E \rightarrow$ Etch the kapton

 $P \rightarrow Plate$ 

#### Wellize

### PEP uRwell examples

Frascati R&D 1D PEP uRwell Active area: 40cm x 5cm



Frascati R&D 1D PEP uRwell Active area: 30cm x 30cm



CLAS12 R&D 2D PEP uRwell Active area: 150cm x 50cm

2023



CLAS12 uRwell rolled in an oven for E-cleaning

# $2016 \rightarrow PicoSec$ precise timing with Micromegas

PICOSEC: Charged particle timing at sub-25 picosecond precision with a Micromegas based detector J. Bortfeldt et. al. (RD51-PICOSEC collaboration), NIM A (903), 2018, <u>https://doi.org/10.1016/j.nima.2018.04.033</u>



Schematic not drawn to scale

X. Wang et al., Study of DLC photocathode for PICOSEC detector, RD51 collaboration meeting, October 2018

DICOSEC

Micromegas

# PicoSec detector modules



**Single pad (2016)** ø1 cm



**Multi pad (2017)** O 1 cm



10x10 module (2022) □ 1 cm

# Other types of MPGD not presented

Rwell → Weizmann :	Rpwell → Weizmann :	High-rate RPCs → INFN & Kobe :
uRwell with THGEM well	RPC with THGEM well	RPC with DLC electrodes
MHSP →Coimbra (Pt) :	Double mesh MM→ USTC China	Micro-PIC→ Kobe :
MSGC with 2 amplification stages		The only dot detector

# MPGD selection Table

Space resolution :	Energy resolution:	Ion back flow:	Rate :	Time resolution :
Micromegas or GEM with capacitive sharing →below 50um	uBULK MM →11% FWHM	multi devices stack → 0.1%	GEMs up to 1Mhz/mm2 resistive MM and uRwell 10Mhz/cm2	PicoSec down to 15ps* GEM 4ns RPCs with DLC*
Cryogenic temperatures:	Low background :	Low mass :	Large size :	Cylindrical detector:
THGEMs , Rwell* , Rpwells*	Micro Bulk MM , uRwell	uRwell, uBulk MM	GEM , floating mesh MM, uRwell	uRwell , GEM and MM (but semi flexible)
	Spherical detectors:	Lower Cost :	Mass production:	
	GEM	uRwell → below 10 000 \$/m2 (material and labor cost included)	uRwell , GEM , MM	*Still in R&D

## Future tasks





-Mass production of vacuum deposited DLC layers

-DLC/Cr/Cu vacuum deposition : needed for high rate Micromegas detectors , resistive GEM and also to simplify uRwell detector production.

-Stronger vacuum deposited photocathodes for PICOSEC detectors.

- -Neutron converter layers : vacuum deposited B4C
- -Transparent conductors : vacuum deposited ITO
- -Other vacuum deposited resistive or converting materials.

#### Conclusion

After millimetric wire Gas chamber

came the Micro Pattern Gas Chambers.

With nanometric resistive and converting layers

we are probably already in the era of NLGCs (Nano layers Gas Chambers).