First Data with CEASERS A new LHC experiment for long lived particle searches on behalf of the FASER collaboration UNIVERSITÉ **DE GENÈVE** at Kruger Conference 2022: Discovery Physics at the LHC

Claire Antel, Geneva University, 6th Dec 2022.





FACULTÉ DES SCIENCES Département de physique nucléaire et corpusculaire









Large Hadron Collider, at HEP energy & intensity frontier.

"The Big 4" titan detectors on ring;

The biggest, CMS & ATLAS, $\sim O(10s)$ meters in diameter & length, have wide physics agenda including searches for new particles

ATLAS & CMS

Focus on central/high pT, EW-scale coupling, since high background + no instrumentation in very forward region









with beam pipe to ATLAS/IP1.

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FASER Overview *Location*

Production at IP1

 $E \sim m_{\pi^0} / \theta_{beam-axis}$

TeV-scale π^0



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A'n



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A'n





FASER Detector Design





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FASER Detector Design





Publications:

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FASER Construction & Commissioning A snapshot

The challenge: Short time scale!





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Dry assembly above surface (end 2020)

Trigger	Rate (Hz)
Veto scintillator station	3.5
Timing scintillator station	10
Pre-shower scintillator station	0.25
Calorimeter	0.25
Random	1
Total rate	15



2021 in tunnel: 125 Million cosmic ray/noise induced events collected

Start of Run 3







Actual FASER detector in actual tunnel.









•FASER (Run 3) sensitive to unprobed phase space for dark photons, ALPs & Neutral Heavy Leptons. •Future FASER 2 (HL-LHC) (assumed 10x bigger) could probe further.

• Sensitivity to New Physics via heavy meson production (e.g. dark Higgs); Larger angle to beam axis.

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FASER Overview



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FASER Overview



- neutrino event shape and kinematic.

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• FASERv (Run 3) to measure ~1000s ν_e , ~10 000s ν_μ and ~10s ν_τ neutrinos in 250 fb^{-1} data.

• To provide: Inclusive charged current neutrino cross-section measurements, constraints on neutrino production in (forward) hadronic interaction models, valuable input to simulation tools related to

_	
	-
VV	
eev	
<u>0-1</u>	
0	
GeV1	



(1/2 proton bunch-crossing precision)





Now well timed in for triggering on charged particles passing through front to back of FASER.



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First Data: Scintillator Performance



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FASER aims to be background free: Need to reject all muons

> 99.995% veto efficiency/layer means negligible muon background after veto: At 1 kHz rate, undetected muons in 4 veto layers $\mathcal{O}(1 \text{ kHz} * 0.00005^{4} \sim 10^{-18}) \text{ Hz}$



First Data: Calorimeter Performance



Good timing measurement in calorimeter important in particular to identify calorimeter-only signals in sync with proton collisions (ALPs->di-photon decay).



First Data: FASERv track reconstruction





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• 3 FASERv emulsion installations so far: 1st box of 0.5 fb^{-1} , 2nd box of 10 fb^{-1} and 3rd box of 30 fb^{-1} exposure.





First Data: FASERv track reconstruction

- Track measurements of first module of 0.5 fb^{-1} showed excellent track resolution.



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3 FASERv emulsion installations so far: 1st box of 0.5 fb^{-1} , 2nd box of 10 fb^{-1} and 3rd box of 30 fb^{-1} exposure.

But higher expected track density in future modules with longer exposure: Maximum track density of $10^6/cm^2$.

First data from 13.6 TeV proton collisions **Reaching top FASER physics rates**



First data from 13.6 TeV proton collisions Reaching top FASER physics rates

• With some unexpected observations....





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First data from 13.6 TeV proton collisions Reaching top FASER physics rates

• With some unexpected observations....

2x higher than predicted rates





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First data from 13.6 TeV proton collisions **Understanding trigger rates**

Rate to proton bunch-crossing dependence paints a picture...



First data from 13.6 TeV proton collisions **Understanding trigger rates**

Rate to proton bunch-crossing dependence paints a picture...





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- ~80% rate correlated with filled bunches (energetic muons)
 - 1/5 leading to coincidence trigger (straight-through muons) •
- ~3% beam 1 background (beam interaction with Q12 leads to background particles, not completely stopped by concrete shielding)
- ~1% cosmics/noise-induced rate
- + ~20% "ambient" rate with no bunch correlation.
 - Suspect related to beam losses (greater at beginning of fill) causing slow moving particle background (neutrons).
- Increased rates not problem for FASER Physics.
 - Trigger system handles extra rate, < 2% deadtime.
 - Beam background can be effectively removed offline with timing cuts.

"ambient" rate with burn off during fill











FASER Data Analysis

Towards calibrating data for physics analyses



Summer 2021 Test Beam Data Analysis

- Dedicated Test Beam at CERN SPS last year to obtain calibration data for calorimeter modules.
- Test Beam data also used to study tracker alignment, data vs simulation, particle identification capabilities.

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Good (~1% at TeV energies) calorimeter resolution achieved. Ongoing studies to understand and improve simulation vs data resolution difference.









FASER is a new Run 3 experiment exploiting full LHC lumi at ~500 m down beamline of ATLAS interaction point, in wait for long lived particles such as dark photons & axions + studying high energy neutrinos.

> Around 40 fb^{-1} of data collected in 2022, with sensitivity to dark photons in unconstrained regions of parameter space.

Interesting things learned about our environment in the tunnel since start of data taking (in particular, unexpected background rates).

Looking ahead

Lots of development for calibration and data vs sim studies ongoing for physics analysis.

Getting ready to provide first physics analysis results for next year's winter conferences.



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Thanks for listening!







The FASER Collaboration

The FASER Collaboration consists of 83 members from 22 institutions and 9 countries

We owe many thanks to:

FASER







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https://faser.web.cern.ch







Back Up



FASER Preshower Upgrade For ALPS->yy

- FASER not capable of separating highly boosted ALP $\rightarrow \gamma \gamma$ decay with current calorimeter: For $E \sim 1$ TeV, < 1 mm photon separations.
- FASER Run 3 upgrade to replace current preshower with high precision monolithic pixel sensors (pixel pitch 100 µm) + tungsten absorber plates (up to 6 radiation lengths).
 - High dynamic range, readout of large number of pixels.
- Lab testing following Test Beam for preproduction ASICs ongoing.





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Charge distribution [fC]



Trigger / pre-shower

station

Magnets

Length: 5 m Aperture: 20 cm Length of decay volume: 1.5 m







Assumptions in Sensitivity

Assumptions by other experiments in sensitivity plots:

NA62 assumes 10¹⁸ protons on target (POT) while running in a beam dump mode that is being considered for LHC Run 3; SeaQuest assumes 1.44×10¹⁸ POT, which could be obtained in two years of parasitic data taking and requires additionally the installation of a calorimeter; the proposed beam dump experiment SHiP assumes $\sim 2 \times 10^{20}$ POT collected in 5 years of operation; the proposed electron fixed-target experiment LDMX during Phase II with a beam energy of 8 GeV and 10¹⁶ electrons on target (EOT); Belle-II and LHCb assume the full expected integrated luminosity of 50 ab⁻¹ and 15 fb^{-1,} respectively; HPS assumes 4 weeks of data at JLab at each of several different beam energies; NA64 corresponds to $5 \times$ 10¹² EOT with 100 GeV energy; and AWAKE is assumed to be working as a fixed-target experiment with a 10-m-long decay volume and 10^{16} EOT accelerated in a 50 – 100 m long plasma cell to the energy O(50 GeV).



Dark Higgs sensitivity Without trilinear coupling



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Dark Photon to visible decays Sensitivity



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Sensitivity to ALPs

Sensitivity to Heavy Neutral Leptons

HNL production mainly through heavy meson and tau lepton decay.

First data from 13.6 TeV proton collisions **Early physics runs dedicated to tracker performance**

Special runs during early collisions used for tracker fine time tuning + hit efficiency scans.

Tracker SCT modules fine time tuned with 390 ps precision. Fine Timing of all stations updated in August.

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Hit efficiency dependent on hit threshold and bias voltage: Expected rise to ~100%, where our nominal values set.

Calorimeter timing for different signals

- Time of arrival of Beam 1 background happens to be perfectly separated from the arrival time of particles from IP1.
- Therefore, Beam 1 background easily reducible even with just calorimeter timing information.

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First Data: FASERv track reconstruction

- Angular distributions of reconstructed tracks in 1st emulsion box:
 - Distribution peaks consistent with particles from the beam line.
 - through 100 m of rock.

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Angular spread of ~0.5 mrad consistent with multiple Coulomb scattering of ~100s GeV muons

FASER Rate vs Collimator Position

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FASER Overview Physics

Large luminosity of highly energetic neutral pions: Source of new boosted physics down beam pipe

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FASER Simulated Signal Efficiency

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mA'=100 MeV, $\epsilon = 10^{-5}$

FASER Background

FLUKA simulation: Magnets very helpfully deflect muons away from beam axis.

Dark photon to visible decay Projections

Ref: <u>Physics Beyond Colliders Working Group Report</u>

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• 10-15 year timescale

