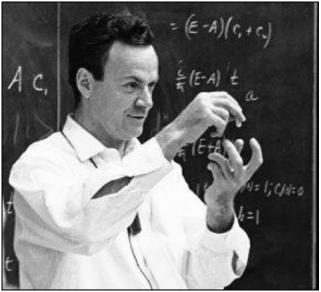




#### Life, the Universe and Everything

"The most remarkable discovery in all of astronomy is that the stars are made of atoms of the same kind as those on the earth."

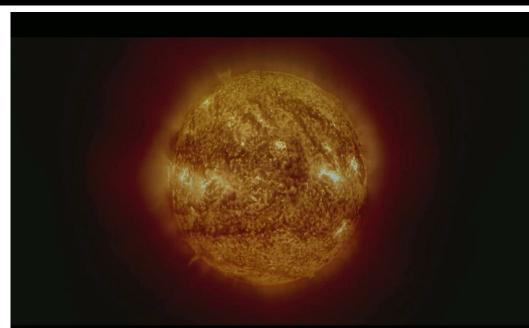
Richard Feynmann Nobel Prize winner





#### Why does the sun shine?

Where does its energy come from?



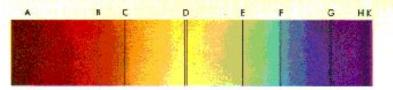
ls it a fire?



Is it a red hot ball of iron?

#### Why does the sun shine?

#### The fingerprint of the stars



THE SOLAR SPECTRUM



SODIUM



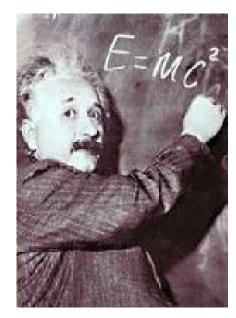
MERCURY



LITHIUM



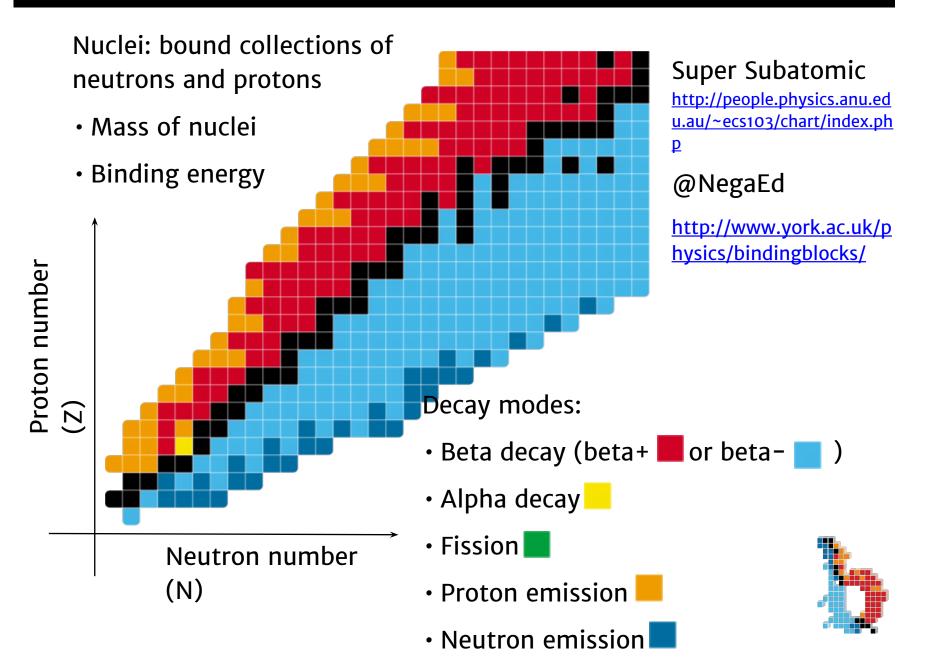
HYDROGEN



 $4 \times H \rightarrow 4He + 2 \times v_e + E$ 

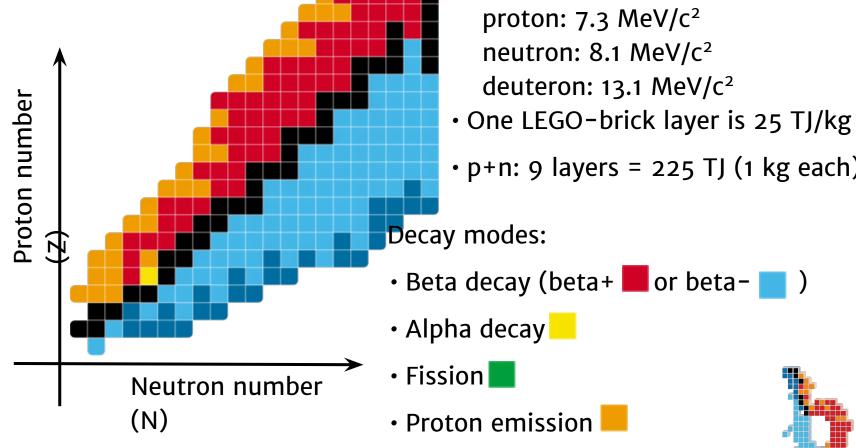


#### Nuclei and decay modes

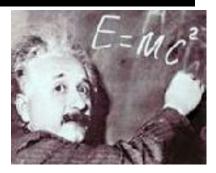


### **Binding Blocks Workshop**

- Decay modes (colours)
- Binding energy per nucleon or mass of nucleus (heights)



Neutron emission



• p+n: 9 layers = 225 TJ (1 kg each)

• Beta decay (beta+ er or beta-

Available energy:

Mass excess is:

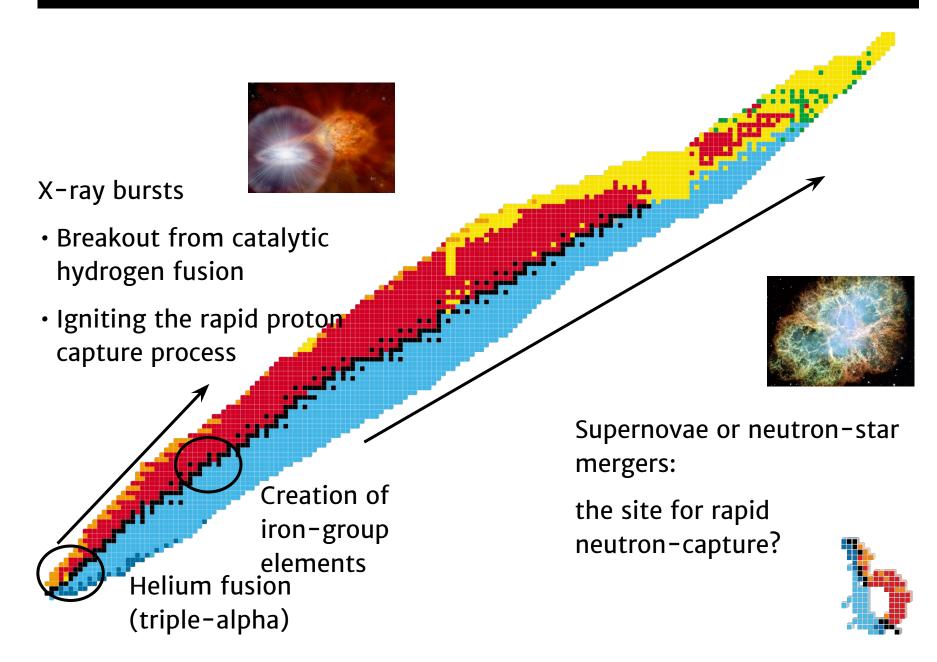


### Binding Blocks Workshop

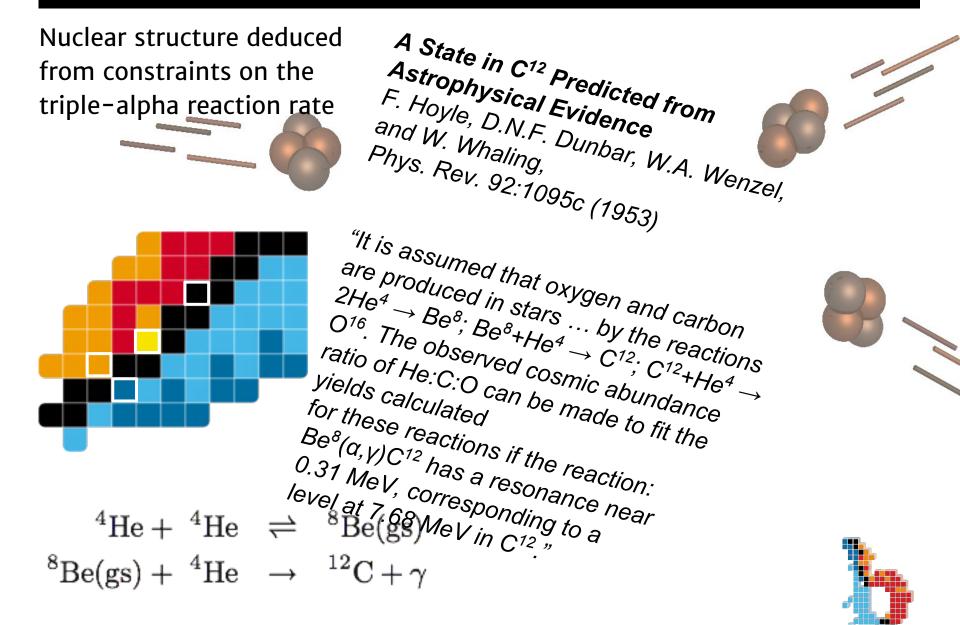
						Ne	<sup>16</sup> Ne 11	<sup>17</sup> Ne 9	<sup>18</sup> Ne 6	<sup>19</sup> Ne 6	<sup>20</sup> Ne 4
					F	<sup>14</sup> F 14	<sup>15</sup> F 9	<sup>16</sup> F 8	<sup>17</sup> F 6	<sup>18</sup> F 5	<sup>19</sup> F 5
				o	<sup>12</sup> O 15	<sup>13</sup> 0 12	<sup>14</sup> 0 7	<sup>16</sup> O 8	<sup>16</sup> O 4	<sup>17</sup> O 5	<sup>18</sup> O 5
			N	<sup>10</sup> N 20	<sup>11</sup> N 14	<sup>12</sup> N 11	<sup>13</sup> N 7	<sup>14</sup> N 6	<sup>15</sup> N 5	<sup>16</sup> N 7	<sup>17</sup> N 7
		с	<sup>8</sup> C 22	<sup>6</sup> C 17	<sup>10</sup> C 11	"C 9	<sup>12</sup> C 5	<sup>13</sup> C 6	<sup>14</sup> C 6	<sup>15</sup> C 8	<sup>16</sup> C 8
		в	<sup>7</sup> B 20	<sup>8</sup> B 16	<sup>9</sup> В 10	<sup>10</sup> B 10	<sup>11</sup> B 8	<sup>12</sup> B 9	<sup>13</sup> B 10	<sup>14</sup> B 12	<sup>15</sup> B 13
		Be	<sup>6</sup> Be 17	<sup>7</sup> Be 14	<sup>8</sup> Be 8	<sup>9</sup> Be 10	<sup>10</sup> Be 10	<sup>11</sup> Be 12	<sup>12</sup> Be 13	<sup>13</sup> Be 15	<sup>14</sup> Be 16
	u	<sup>4</sup> Li 29	<sup>6</sup> Li 14	<sup>6</sup> Li 14	<sup>7</sup> Li 13	<sup>8</sup> Li 15	<sup>9</sup> Li 16	<sup>10</sup> Li 18	<sup>11</sup> Li 19		
	He	<sup>3</sup> He 24	<sup>4</sup> He 7	<sup>5</sup> He 14	<sup>6</sup> He 16	<sup>7</sup> He 19	<sup>е</sup> Не 20	<sup>9</sup> He 23	<sup>10</sup> He 24		
н	1H 33	²H 30	<sup>3</sup> H 24	⁴H 29	<sup>5</sup> Н 30	°H 32	'H 32				
	(n)	<sup>1</sup> n 36									

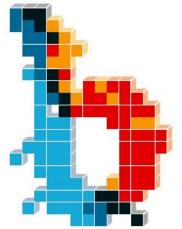


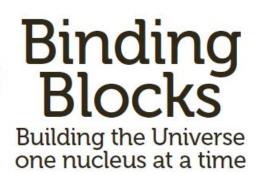
#### Nucleosynthesis in stellar explosions



#### Nucleosynthesis in stellar explosions







UNIVERSITY of rk



# Nuclear fusion

Hannah Willett, York Plasma Institute

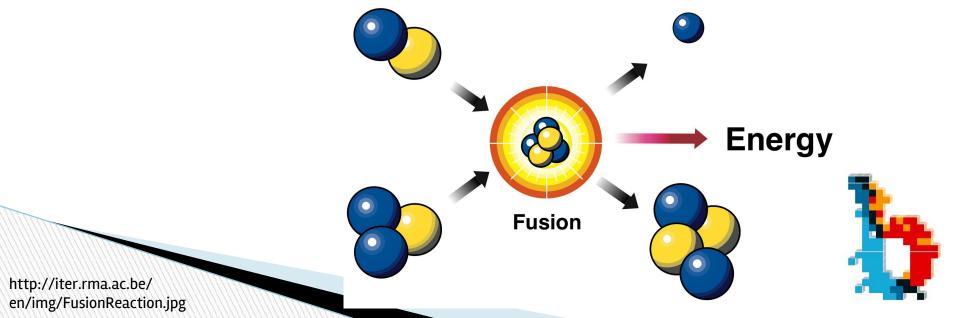
### A clean energy source?

- Fusion is a massive energy source powers stars (more on that later)
- Fossil fuels are finite resources, plus climate change...
- Fusion fuels come from seawater and lithium (we have lots of both!)
- Can we harness fusion in a power plant?
  - "Star in a jar"

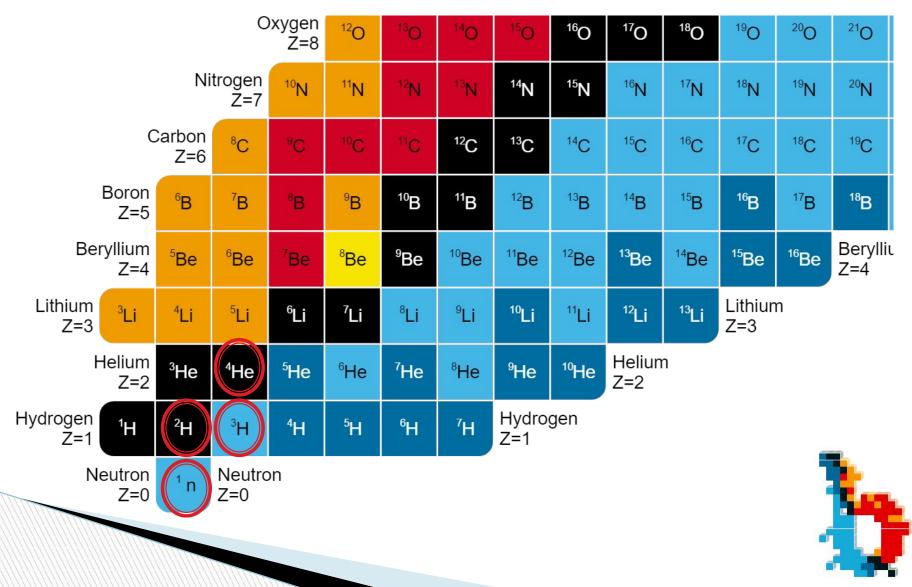


# What is fusion?

- Sticking two small nuclei together to make a larger one
- Stars use hydrogen (protons) as fuel
- Use deuterium and tritium on Earth
  - More achievable temperatures and pressures



# **DT** fusion



### How much energy?

- 34 layers of bricks is ~850000 GJ
- VK energy use per capita (2012) ~125 GJ
  - Electricity use per capita (2011) ~20GJ
- Half a bathtub of seawater and the lithium in a laptop battery would provide ~30 years of *electricity* use for one person
- Fusion could be the energy source of the future...

