Calorimetry

Introduction and outlook on future projects

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1) Interactions of particle with matter and basics of electromagnetic showers

2) Interactions of hadrons and hadronic showers

3) Particle detection

4) Calorimeter concepts (not only) for future projects ...



R. Wigmans: Calorimetry

D. Wegener: Detektoren in der Teilchenphysik - Lecture Uni Dortmund
W.R. Leo: Techniques for Nuclear and Particle Phyisics Experiment
C. Grupen: Teilchendetektoren
Sitar et al.: Ionization measuremens in High Energy Physics
arxiv:1412.2653, 1507.05893, 1602.08578

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Lots of Material 'stolen' from presentations and articles found on the Web.

If you find that I have used your material without a citation please write me and I will include the reference

Thanks to Hengne Li for producing several figures for this lecture



Chapter 1

Interactions of particles with matter and basics of electromagnetic showers



4π 'Germanium Ball' of AGATA Experiment



ATLAS TileCal Barrel Calorimeter



1m



Calorimeters are employed in 'table top' experiments and in huge experimental apparatuus

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ATLAS Calorimeter



- em Barrel : (|η|<1.475) [Pb-LAr]
- em End-caps : 1.4<|η|<3.2 [Pb-LAr]
- Hadronic End-cap: $1.5 < |\eta| < 3.2$ [Cu-LAr]
- Forward Calorimeter: 3.2<|η|<4.9 [Cu,W-LAr]</p>
- ~190K readout channels
- Hadronic Barrel: Scintillating Tile/Fe calorimeter





Scattering Center: Nucleus or Atomic Shell

Detection Process is based on Scattering of particles while passing detector material

Energy loss of incoming particle: $\Delta E = p_0 - p'_0$



'Real' Bethe Bloch Formula

After consistent quantum mechanical calculation Valid for particles with $m_0 >> m_e$

 ΔE = 0: Rutherford Scattering

 $\Delta E \neq 0$: Leads to Bethe-Bloch Formula

$$\frac{dE}{dx} = 4\pi N_A r_e^2 m_e c^2 z^2 \frac{Z}{A} \frac{1}{\beta} \left[\frac{1}{2} \ln \left(\frac{2m_e c^2 \gamma^2 \beta^2 T_{\text{max}}}{I^2} \right) - \beta^2 - \frac{\delta}{2} \right]$$

$$\beta$$
, γ - Velocity, gamma-factor of incoming particle

- Z, A Nuclear charge and mass of absorber
- r_e, m_e Classical electron radius and electron mass
- N_A Avogadro's Number = 6.022x10²³ Mol⁻¹
 - Ionisation Constant, characterizes Material typical values 15 eV
- δ Fermi's density correction
- T_{max} maximal transferrable energy iThemba School 2023 - Calorimeters Chapter 1



Minimal Ionizing Particles (MIPS)

dE/dx passes Broad minimum @ $\beta \gamma \approx 4$

Contributions from energy losses start to dominate kinematic dependency of cross sections

1.13

1.51

1.82

Lead

Steel

02

Power [MeV cm²/g] µ⁻ on Cu 100 Radiative Anderson Ziegler Bethe Bloch 10 Lindhard Scharff Ε. Radiative Losses Minimum ionisation Stopping Nuclear Losses Without & 0.01 104 105 106 0.001 0.1 10 100 1000 βγ 10 10 166 | 16 166 166 | 11 11 B.1 typical values in Minimum [MeV/c] [GeV/c] [TeV/c] Muon Momentum [MeV/(g/cm²)] [MeV/cm] 20.66 **Role of Minimal Ionizing Particles ?** 11.65

 $2.6 \cdot 10^{-3}$

Discussion of Bethe Bloch Formula V

u-___

Anderson-

Ziegler

0.01

0.1

10

[MeV/c]

1

100

Radiative losses - Not included in Bethe-Bloch Formula

Particles interact with Coulomb Field Of <u>nuclei</u> of Absorber atoms

Stopping Power [MeV cm²/g]

100

01 Lindhard-Scharff

0.001

0.1

Nuclear Losses



Energy loss due to Bremsstrahlung

Important for e.g. muons with E > 100 GeV Dominant energy loss process for electrons (and positrons) $(m\mu/m_e)^2 \sim 40000$

Muon Momentum

100

10

[GeV/c]

μ[−] on Cu

Bethe Bloch

10

1

βγ

Radiative

104

1

1000

100

Radiative Losses

Without δ

105

10

[TeV/c]

106

100



Critical Energy

Point where $\Delta E_{lon} = \Delta E_{Brems}$ 200 For materials e⁻ in Cu Copper 610 MeV $X_0 = 12.86 \text{ g cm}^{-2}$ (from PDG) in solid and liquid \mathcal{E}_{c} $E_{c} = 19.63 \text{ MeV}$ 100 Z + 1.24phase $dE/dx \times X_0 (MeV)$ 70 Erac monstran Rossi: 50 Ionization per X_0 Brons $\frac{710 \text{ MeV}}{Z + 0.92}$ 40 = electron energy For gases \mathcal{E}_{c} 30 Ionization 20 **Emperical values** Brems = ionization Based on fits to data 10 5 10 20 2 50 100 200 Electron energy (MeV) $\epsilon_{\rm c}$ is a characteristic parameter of material e.g. ϵ_c for Uranium 6.75 MeV



General: Beer's Absorption Law: $I = I_0 e^{-\mu x}$, $\mu = Absorption coefficient^{I}$



Different processes for different energies:

- 1) Photoelectric Effect E ≤ 500 keV
- 2) Compton Scattering 500 keV < E < 5 MeV
- 3) Pair Production E > 2m_e



Above all relevant electromagnetic interactions were introduced

For electrons/positrons

a) Ionizationb) Bremsstrahlung(Annihilation for Positrons)

For Photons

- a) Photoelectric Effect
- b) Compton Scattering
- c) Pair Production
- A primary particle looses energy due to one of these interactions in a given absorber material
- The produced secondaries in turn do interact with the material and produce themselves further particles

Development of an electromagnetic cascade



High energetic particles

Energy loss of electron by Bremsstrahlung: Photons convert into e⁺e⁻-Pairs



For high energetic particles shower contain thousands of particles

$$\frac{dE}{dx} = -\frac{E_e}{X_0} \Longrightarrow E = E_0 e^{-x/X_0}$$

Energy loss after X₀: E₁ =E₀/2
 Photons -> materialize after X₀

 $E_{\pm}=E_1/2$

Number of particles after t: $N(t) = 2^t$ Each Particle has energy

$$E = \frac{E_0}{N(t)} = E_0 2^{-t} \Longrightarrow t = \ln\left(\frac{E_0}{E}\right) / \ln 2$$

Shower continues until particles reach critical energy (see p. 16) where $t_{max} = ln(E_0/\epsilon_c)$

Shower Maximum increases logarithmically with Energy of primary particle (Important for detector design !!!)

Scaling variables I – Radiation Length

 $\Phi_{rad} = \frac{1}{E_0} N \int_{0}^{E_{\gamma 0}} E_{\gamma} \frac{d\sigma_{rad}}{dE_0} dE_{\gamma} \Longrightarrow - \frac{dE}{dx} = N E_0 \Phi_{rad}$

 $-\left(\frac{dE}{dx}\right)_{rad} = N \int_{0}^{E_{\gamma 0}} E_{\gamma} \frac{d\sigma_{rad}}{dE_{\gamma}} dE_{\gamma}$

 $\frac{d\sigma_{rad}}{dE_{\gamma}} \sim \frac{1}{E_{\gamma}} f(Z)$

(Not only) For Bremsstrahlung:

Spectrum of radiated photons:

 $\Phi_{\rm rad}$ is independent of the energy of the radiated photon and only a function of the the material

Radiation Length X₀: Distance after which a particle has lost 1/e due to radiation

$$E = E_0 \exp\left(-\frac{x}{X_0}\right) \quad \text{With } X_0 = 1/N\Phi_{\text{rad}} \Rightarrow X_0 = \frac{A}{4\alpha N_A Z(Z+1)r_e^2 \ln(183Z^{-1/3})} [g/cm^2]$$

Absorption Coefficient Photons

$$\mu(E >> m_e c^2) = \frac{28}{9} n Z^2 \alpha r_e^2 \ln \frac{183}{Z^{1/3}} = \frac{7}{9} X_0^{-1}$$

i.e. Photons travel a longer distance before they interact iThemba School 2023 - Calorimeters Chapter 1

Some values: $X_{0,Air}$ =30 420 cm $X_{0,AI}$ = 8.9 cm $X_{0,Pb}$ = 0.56 cm



Longitudinal shower profile



Shape of shower profiles Material independent = Energy loss Independent of material

 \dots as function of X_0

Differences:

- Pair production by emitted photons increases with Z and extends to much lower energies
- Critical energy is Z-dependent e.g. 43 MeV for AI 7 MeV for Pb
- X₀ looses meaning at low energies

Scaling variables II – Molière Radius

2 Effects 1) Electrons (and positrons) undergo <u>multiple scattering</u> in the Coulomb field of absorber nuclei $\Theta_{mean} = 0$ with standard deviation Primary e[±] $\Theta_{rms} = \sqrt{\langle \Theta \rangle} = \frac{13.6}{\beta cp} \sqrt{\frac{x}{X_0}}$ For projected Scattering angle distribution

2) Isotropic production of secondaries by Photoelectric effect and Compton Scattering

Low energetic component of shower Expect different behaviour in region 'far' away from shower axis

Transversal extension of shower characterized by Molière Radius

$2 \sqrt{1 - X_0}$		Typical Values for ϵ_c , ρ_M :		
$\rho_{M} = mc^{2}\sqrt{4\pi} / \alpha - \frac{0}{2} = 21.2 MeV - \frac{0}{2}$		ϵ_{c} [MeV]	R _M [cm]	
\mathcal{E}_{C} \mathcal{E}_{c} \mathcal{E}_{c}	Pb	7.2	1.6	
90% of shower energy is contained within p_M	NaJ	12.5	4.4	
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Energy resolution – The master formula



$$\frac{\sigma(E)}{E} \propto \frac{\sigma(N)}{N} \propto \frac{1}{\sqrt{N}} \propto \frac{1}{\sqrt{E_0}}$$

holds also for hadron calorimeters

Also spatial and angular resolution scale like $1/\sqrt{E}$

Relative energy resolution of a calorimeter improves with E_{0}

More general:





Backup



Energy loss due to collisions

à la Leo: Partially taken from lecture at NIKHEF by unknown author



Energy loss due to collisions cont'd

$$\Delta p_{\perp} = \int F_{\perp} dt = \int_{-\infty}^{\infty} \frac{Z_{2} e \gamma Z_{1} e dt}{\left(b^{2} + \gamma^{2} v_{1}^{2} t^{2}\right)^{\frac{3}{2}}}$$
$$\Delta p = \Delta p_{\perp} = Z_{2} e \gamma Z_{1} e \left[\frac{t}{b^{2} \left(b^{2} + \gamma^{2} v_{1}^{2} t^{2}\right)^{\frac{1}{2}}}\right]_{-\infty}^{\infty} = \frac{2Z_{1} Z_{2} e^{2}}{v_{1} b}$$

$$\Delta E = \frac{(\Delta p)^2}{2m} = \frac{2Z_1^2 Z_2^2 e^4}{b^2 v_1^2 m}$$

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des 2 Infinis

(Particle 2 is in rest -> non-relativistic calculation of energy loss possible)

Interactions with electron: $m = m_e, Z_2 = 1$ Interactions with nucleus with mass number A and atomic number Z: $m = Am_p \approx 2Zm_p, Z_2 = Z$

$$\frac{\Delta E(e)}{\Delta E(nucleus)} = \frac{2Z_1^2 e^4}{b^2 v_1^2 m_e} \left/ \frac{2Z_1^2 Z^2 e^4}{b^2 v_1^2 2Z m_p} = \frac{2m_p}{Z m_e} \approx 4000 / Z \right.$$

-> Energy loss due to collisions is dominated by interactions with the electrons (NB: we are comparing interactions with 1 electron to interactions with the nucleus of an atom, a non-ionized atom has Z electrons) a School 2023 - Calorimeters Chapter 1 21



So far interactions with single electrons but material consists of many electrons and nuclei

N. Bohr: Particle passes through center of thin shell



Value for b_{max}



Determination of energy loss

Transversal Field Strength: $(b^2 + \gamma^2 v_1^2 t^2)^{3/2}$

'Interaction' Time:

 $b/v_1\gamma$

Interaction Time < Orbital Frequency ω of Electrons such that binding effects can be neglected

 $\Rightarrow b_{max} = \gamma v_1 / \omega$

Maximal Energy Transfer for m >> m_e (see above): $2m_e\beta^2\gamma^2$

Using:
$$\Delta E = \frac{2Z_1^2 e^4}{b^2 v_1^2 m_e}$$
 it follows: $b_{\min} = \frac{Z_1 e^4}{\gamma v_1^2 m_e}$ so:
$$\frac{dE}{dx} = \frac{4\pi n_e Z_1^2 e^4}{m_e v_1^2} \ln \frac{m_e v_1^3 \gamma^2}{Z_1 e^2 \omega}$$

Detailed discussion follows





Drop continues until $\beta \gamma \sim 4$

Muon Momentum

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CLab Discussion of Bethe Bloch Formula IV



Remark: Density Correction tightly coupled to Čerenkov-Effect

P does not contribute to

energy loss !



v ~ c, i.e. electrons and positrons

See e.g. E.A. Uehling 1954, Ann. Nucl. Sci. 4, 315 Sect. 1.1 Results for Electrons:

$$\frac{dE}{dx} = 4\pi N_A r_e^2 m_e c^2 z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[\frac{1}{2} \ln \left(\frac{2m_e c^2 \gamma^{3/2}}{\sqrt{2}I^2} \right) + \frac{1}{16} - \frac{\delta}{2} \right]$$

Results for Positrons:

$$\frac{dE}{dx} = 4\pi N_A r_e^2 m_e c^2 z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[\frac{1}{2} \ln \left(\frac{2\sqrt{2m_e} c^2 \gamma^{3/2}}{I^2} \right) + \frac{23}{24} - \frac{\delta}{2} \right]$$

These particles are already ultrarelativistic at E ≈100 MeV