

Exploring the α -Process with High Precision (p,t) Reactions

Type 1 X-ray Bursts are identified as thermonuclear explosions on the surface of accreting neutron stars. These bursts can be observed and characterized through their bolometric luminosity curves. In a small number of type 1 X-ray bursts, a double peak structure in the light curve has been observed. This double peak structure lead Fisker et al to propose a nuclear waiting point impedance in the thermonuclear reaction flow of the α -process, and conclude that uncertainties in (α,p) reactions rates on these potential waiting points can be directly observed in the structure of X-ray burst light curves [1]. (α,p) reaction rates on two possible waiting point nuclei, ^{26}Si and ^{34}Ar , were examined by investigating the level structure in the respective compound nucleus, ^{30}S and ^{38}Ca . This was done through high-precision measurements of $^{32}\text{S}(p,t)^{30}\text{S}$ and $^{40}\text{Ca}(p,t)^{38}\text{Ca}$ reactions utilizing the high energy-resolution zero-degree techniques with the K600 spectrometer at iThemba LABS [2]. States above the α -threshold have been identified and precise excitation energies were determined. These excitation energies, along with calculated level parameters were then used to determine the reaction rates for the $^{26}\text{Si}(\alpha,p)$ and $^{34}\text{Ar}(\alpha,p)$ reaction over stellar temperature ranges. Experimental excitation energies along with calculated reactions rates will be presented.

Ref.

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[2] R Neveling, H Fujita, F D Smit, T Adachi, G P A Berg, E Z Buthelezi, J Carter, J L Conradie, M Couder, R W Fearick, S V Fortsch, D T Fourie, Y Fujita, J Gorres, K Hatanaka, M Jingo, A M Krumbholz, C O Kureba, J P Mira, S H T Murray, P von Neumann-Cosel, S O'Brien, P Papka, I Poltoratska, A Richter, E Sideras-Haddad, J A Swartz, A Tamii, I T Usman, and J J van Zyl. High energy-resolution zero-degree facility for light-ion scattering and reactions at iThemba LABS. *Nuclear Inst. and Methods in Physics Research, A*, 654(1):29{39, October 2011.

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