Pygmy resonances in the mirror

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Pygmy Dipole Resonance	• Electric dipole transition strength on the lower-energy tail of the giant dipole resonance	
Origin and implications	 Isospin asymmetry? Neutron/proton skin? Shell structure? Deformation? Clustering? Symmetry energy and capture reactions 	
This work	 Compare a chain of loosely bound isotones (N=20) to their stable mirrors (Ca isotopes) The absolute asymmetry is the same; the pygmy resonance is not. 	
Conclusion	 Coulomb field → loose binding: hugely important Prediction for a coherent proton-pygmy resonance in ⁴⁶Fe - and ⁴⁴Cr? 	and further explorations of symmetry energy and effective mass under way – with KIDS* *See talk

What is called a pygmy dipole resonance?

- When the photonic field acts on the atomic nucleus, the giant dipole resonance (GDR) is strongly exited. In many heavy nuclei it has a smooth shape, which can be described by one or two Lorenzian curves.
- Depending on the nucleus, there is often excess transition strength at lower energies, which does not fit with the GDR: D.M. Rossi et al.
- That excess strength may or may not appear as a clean resonance peak. But it is customarily called a pygmy dipole resonance (PDR) until further notice.

What is it then?

- The PDR is observed in asymmetric nuclei (N>Z). Asymmetry must play a role - but how?
- An exotic mechanism proposed early on is the neutron-skin mode: An asymmetric nucleus develops a neutron skin; that skin can oscillate against the isospin-symmetric core.
- It can be an unglamorous shell effect: Single-particle transitions across one $\hbar\omega$

Let us look in the mirror then: N=20 vs Z=20

Ni-48 vs Ca-48

Below, "CRPA" means a complete treatment of particle states, i.e., continuum; while in "QRPA" a harmonic-oscillator basis is used.



Clues: the least bound occupied state, $f_{7/2}$



- and proton skin thickness in Ni-48 vs neutron skin thickness in Ca-48: 0.25 vs 0.15 fm
 - Whatever happened to isospin symmetry?

shell that "failed" to join the GDR.

Or both in some degree or other – or something else yet.

Why do we care?

- If there is a vibrational mechanism involved (neutron-skin oscillation?), there must be a restoring force, plausibly related to the symmetry energy. We want to know that - for nuclear structure and astrophysical modeling.
- The PDR is typically observed in the vicinity of the particle emission threshold. Resonances near a threshold can influence capture reaction rates drastically. We must factor them in when simulating nucleosynthesis paths.
- As a shell effect it can unlock information on nuclear structure.

What are the issues?

- Where to begin. Is it in fact vibrational? Enhanced in r-process nuclei? What is the role of asymmetry? What is the role of deformation? ...
- ... Model validation: What is the role of the continuum? What is the role of phonon coupling? Of clustering?

The strong violation of symmetry shows the importance of the Coulomb field. Owing to the Coulomb repulsion there are very loosely bound protons in Ni-48. A small binding energy means a very extended wave function, hence the above results. The Coulomb barrier is no remedy.



What seems to be rather symmetric is the collective isoscalar state, IS-LED – see transition densities on the left. But it would be difficult to make out experimentally, being embedded in a continuum of mixed-isospin states.

All N=20, A=40-48 isotones (QRPA)

- Similar conlusions
- Fe-46, Ni-48: > 1% TRK Any sign of coherence? Proton-skin mode?



A microscopic analysis points to Fe-46 as a strong candidate for a coherent isovector state, i.e., a possible proton-skin oscillation; in Cr-44 also likely.

And what about the equation of state,

The PDR is likely of mixed isospin: it responds (partly?) to isoscalar probes. Then what is the role of the collective isoscalar mode ("IS-LED"), in the same energy region, which has nothing to do with asymmetry and skins?

There is more than the PDR at the bottom of the GDR curve... [PP, EXON2016]

•If asymmetry plays a role, what about *absolute* asymmetry? Is N>Z the same as Z<N? Why not?

Present work: mirror (a)symmetry

Based on:

Y. Kim and P.P., Eur. Phys. J A 52, 176 (2016).

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the nucleon effective mass m*, etc?

Here are some more model results (left). The symmetry-energy slope parameter L may play a role; m^{*} may have an effect insofar as it affects the Fermi energy and shell spacing; all else equal, the m^{*} effect is like this: \downarrow



S.Krewald and J.Speth, Int. J. Mod. Phys. E 18, 1425 (2009) D. Savran et al., Prog. Part. Nucl. Phys. 70, 210 (2013)