

Evolution to neutron rich isotopes in the fine structure of the Isoscalar Giant Monopole Resonance in $^{40,42,44,48}\text{Ca}$ using alpha inelastic scattering at zero degrees

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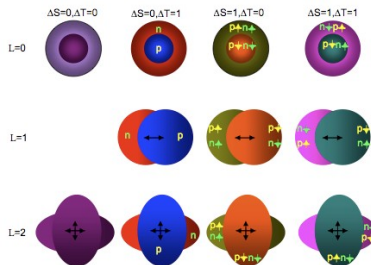
Prof. J. Carter (Wits University)

COMEX6, Nov. 1, 2018



What are giant resonances?

- GRs are the collective motion of the nucleons within the nucleus
- Resonance typically occurs at excitation energies between 10 - 30 MeV
- Resonance can be **isoscalar** (i.e. the nucleons move in phase) or **isovector** (i.e. the nucleons move out of phase)
- The nucleus can take on **monopole** ($L = 0$), **dipole** ($L = 1$), **quadrupole** ($L = 2$), etc, structures
- ISGMR $\rightarrow \Delta L = 0, \Delta S = 0, \Delta T = 0$



Motivation

Fine structure

- GRs carry fine structure and experimental study of the GRs provides insight into the main mechanism leading to the fine structure phenomenon
- Fine structure observed in GRs:
 - ISGQR → coupling of elementary 1p-1h states to low-energy surface vibrations
 - IVGDR → fragmentation of 1p-1h strength (Landau damping)
 - M1, M2, GTR → Mixing of 1p-1h with np-nh states
 - ISGMR → ?
- Detailed study of the excitation and decay of ISGMR is imperative



Motivation

GRs and Nuclear incompressibility

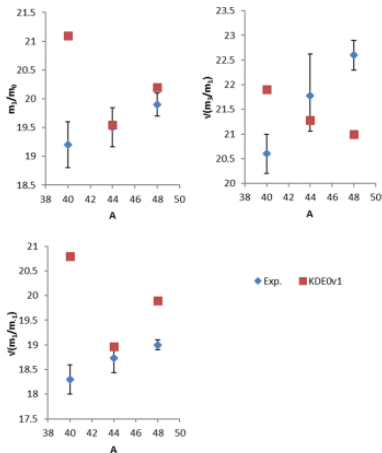
- Giant resonances are important because they provide information about bulk properties of the nucleus e.g. compression modulus.
- Excitation energy of GMR is directly related to the incompressibility of the nucleus, an important parameter in the EOS of nuclear matter.

$$E_{ISGMR} = \frac{\hbar\sqrt{K_A}}{m\langle r^2 \rangle} \quad (1)$$

where E_{ISGMR} is the centroid energy of the GMR, m is the mass of the nucleus, $\langle r^2 \rangle$ is the gs mean-square radius, and K_A is the incompressibility of the nucleus



- GR energies generally go roughly as $A^{-1/3}$.
- Contrary to expectations, J. Button *et al.* (PRC 96, 054330 (2017)) reported ISGMR energy in ^{48}Ca higher than in ^{40}Ca .



- It is imperative to examine this anomaly.



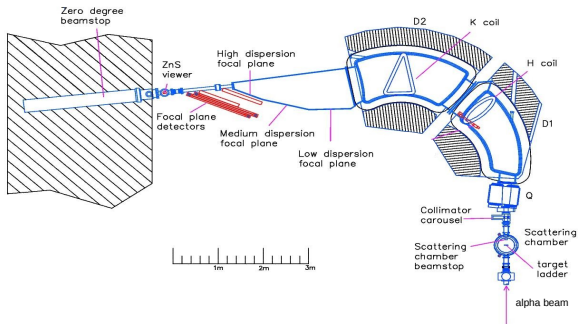
Objectives

- Obtain completely background-free inelastic scattering spectra for the $^{40,42,44,48}\text{Ca}$ isotopes
- Investigate the fine structure of the ISGMR as a function of neutron excess in the Ca isotopes and extract its characteristics energy scales
- Extract the $J^\pi = 0^+$ level densities
- Compare the value of the nuclear incompressibility K_A , based on the ISGMR results, with existing values



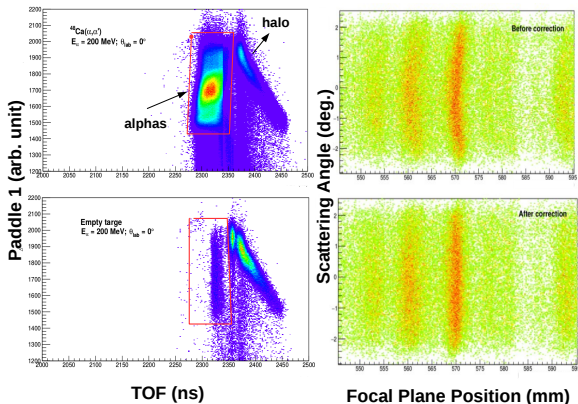
Measurements with the K600 Magnetic Spectrometer at iThemba LABS

- Reaction: inelastic scattering of 200 MeV α particles at 0° and 4°
- ISGMR partly overlaps with ISGQR, hence, 4° measurements were taken to remove ISGQR contribution from the 0° measurements
- Targets: ^{24}Mg (2.1 mg/cm^2) and $^{42,44,48}\text{Ca}$ ($1.35, 1.43, 1.09 \text{ mg/cm}^2$)



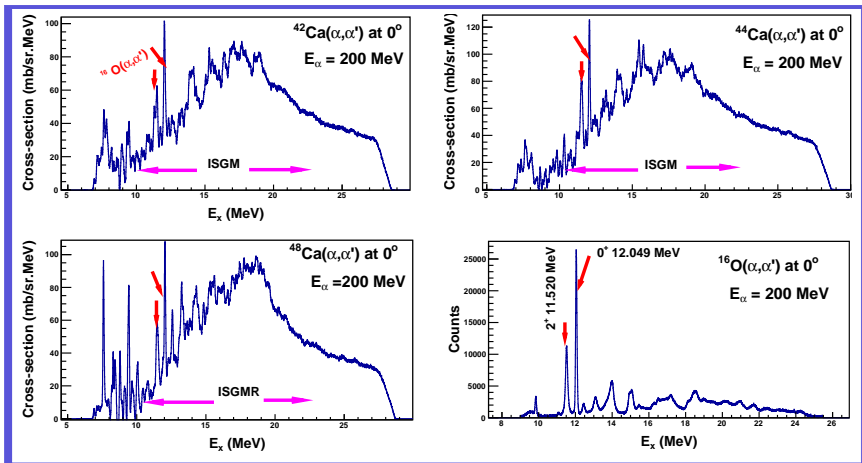
Background reduction and software correction

- Particle ID (PID) and time-of-flight (TOF) cuts were made to select alphas AND reduce background contributions from small angle scattering off the collimator lips.
- Lineshape correction was done to optimise the resolution.



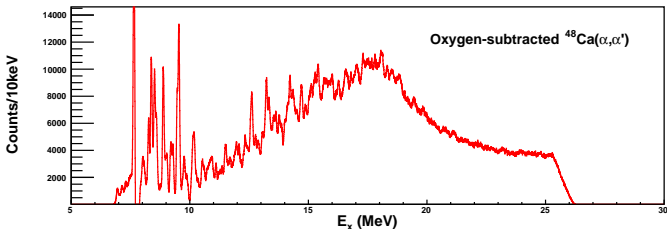
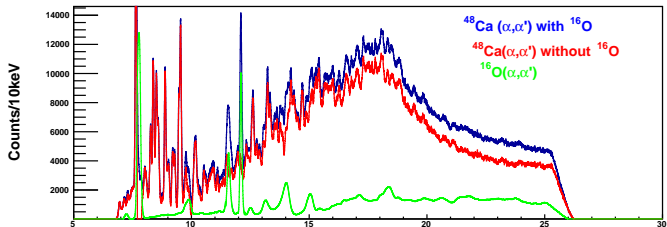
Excitation energies of Ca isotopes & ^{16}O

- The ^{16}O contamination peaks found on the lower edge of the ISGMR region

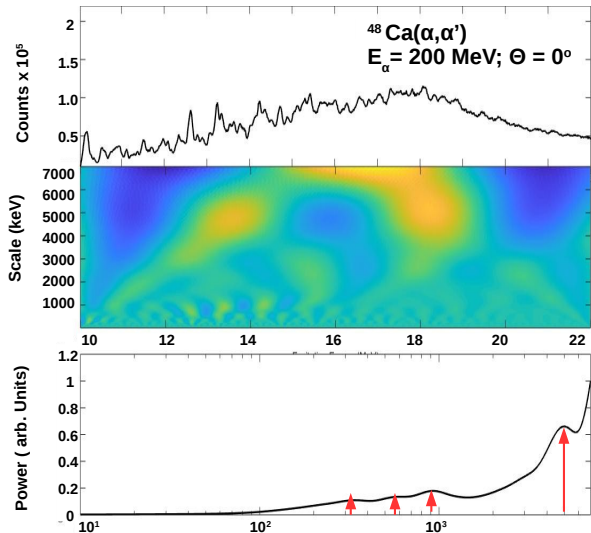


Removal of ^{16}O contamination

- The ^{16}O contamination was removed through a direct background subtraction.



Wavelet Analysis on ^{48}Ca without ^{16}O



Comparison of scales

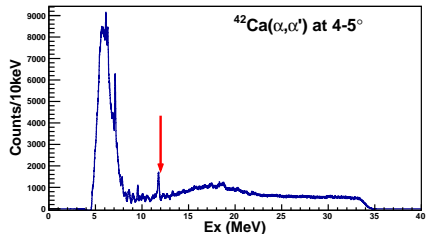
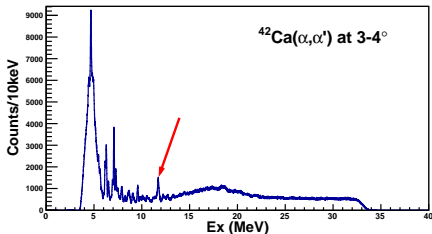
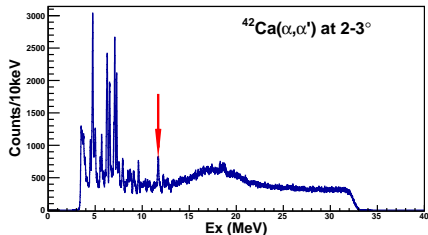
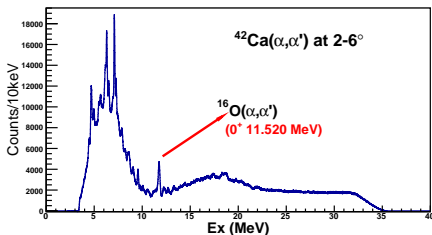
The excitation energy range considered is 10 – 22 MeV

Target	Class I $\Delta E \leq 300$ (keV)	Class II $300 < \Delta E < 1000$ (keV)		Class III $1000 < \Delta E < 6000$ (keV)	
^{16}O	- -	540	930	2390	5210
^{48}Ca (with ^{16}O)	290	560	960	3000	4970
^{48}Ca (without ^{16}O)	- -	340	610 930	-	4990

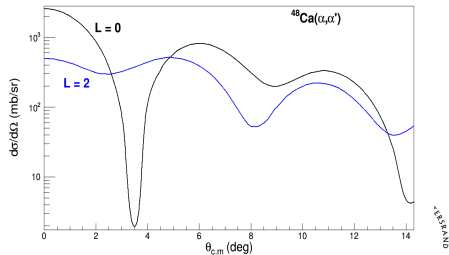
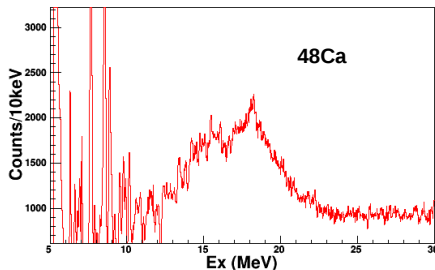
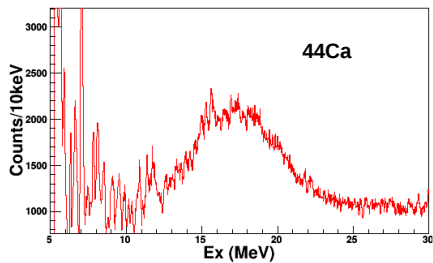
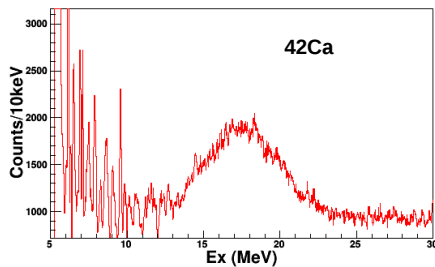
The removal of the ^{16}O has no significant effect on the structure of the target of interest.



- ^{16}O contamination at around $Ex = 11.5$ MeV.

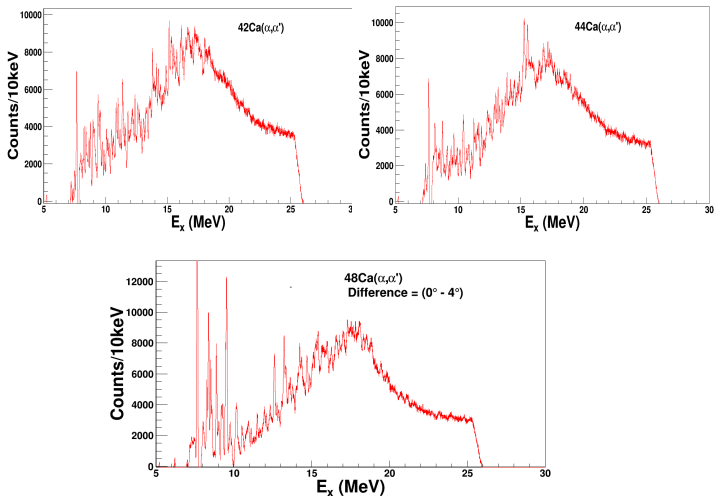


- DWBA plot: 1st minimum of ISGMR at 3-4°



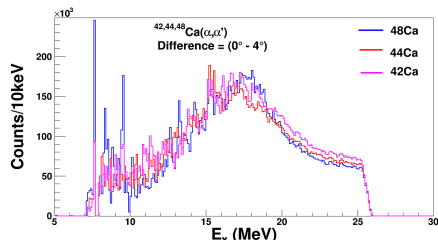
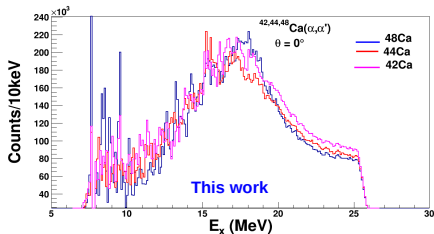
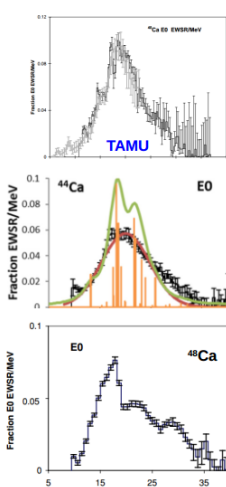
Y. ERBAUD

- Difference-spectrum is essentially monopole monopole



Structure effects in Ca isotopes

- TAMU recommends the need to consider nuclear structure effects in order to understand the systematics of Ca isotopes.



- Continuous Wavelet Analysis on the (α, α') data to extract characteristics energy scales from fine structure of ISGMR
- Comparison of scales from experimental data with those from theoretical calculations
- Autocorrelation function analysis to extract $J^\pi = 0^+$ level densities



Collaborators...

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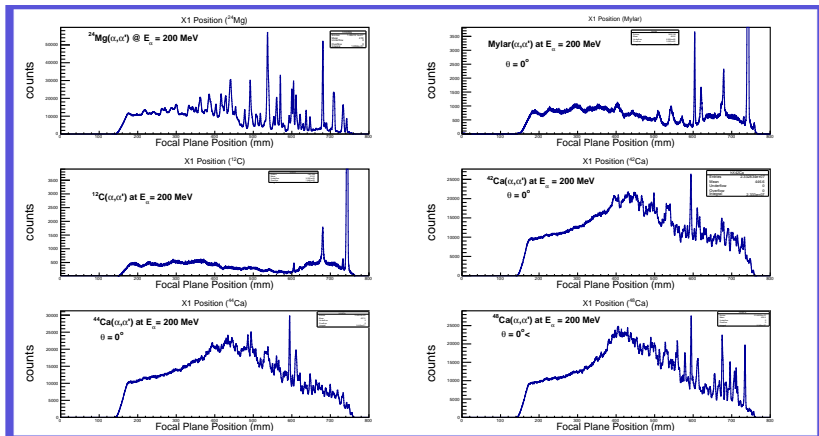


Backup Slides

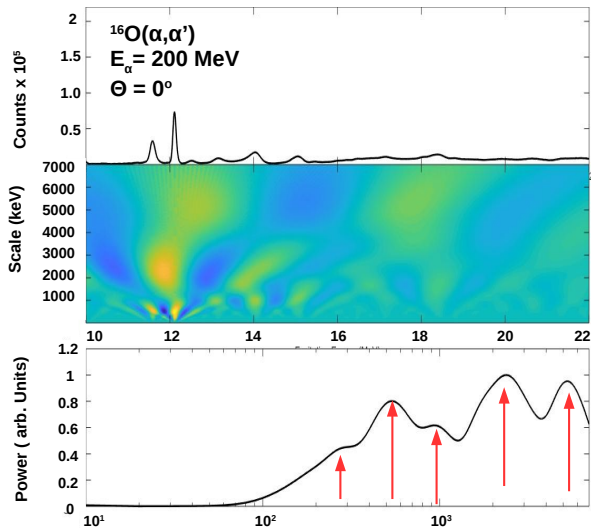


Position spectra

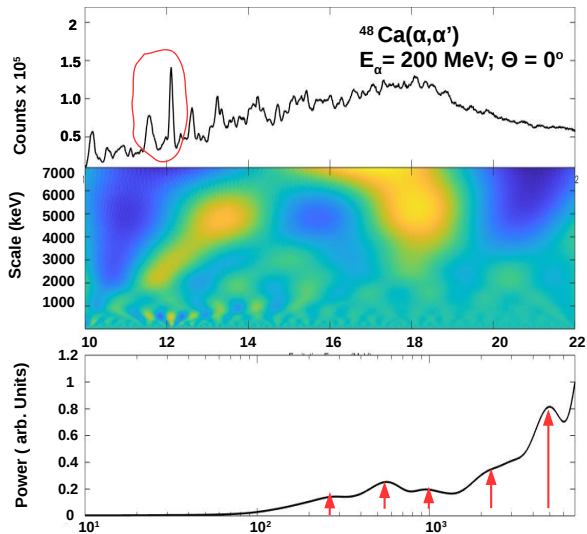
- Excitation energy calibration was done using the $^{24}\text{Mg}(\alpha, \alpha')$, whose spectrum exhibited well-known excited states
- Mylar target was used to check/reveal the presence of structures due to ^{16}O contamination



Wavelet Analysis on ^{16}O

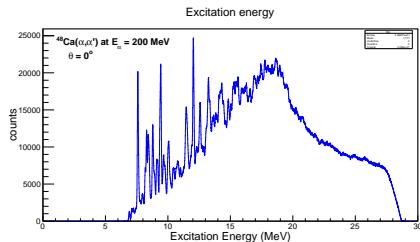
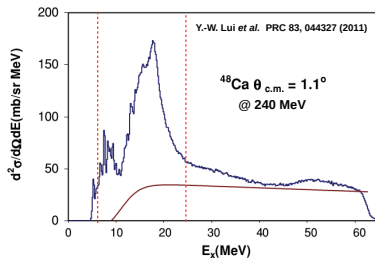


Wavelet Analysis on ^{48}Ca with ^{16}O



Comparison with previous measurements

- Texas A&M has data for only $^{40,48}\text{Ca}$ isotopes while this work presents data for $^{40,42,44,48}\text{Ca}$ with better resolution
- Texas A&M data on Ca isotopes does not allow for fine structure analysis



Example

$$K_A = K_V + K_{surf} A^{-1/3} + K_{sym} \left(\frac{N - Z}{A} \right)^2 + K_{Coul} \frac{Z^2}{A^{4/3}} + \dots$$

Volume term, $K_V = K_{(A \rightarrow \infty)} = K_{nm}$, where K_{nm} compressibility of the nuclear matter. Let $\delta = (N - Z)/A$

- In recent years, studies of the symmetry term K_{sym} are mostly on heavy nuclei, especially on Sn isotopes, yielding varying results (T. Li *et al.*, Phys. Rev. C 81, 034309 (2010)).

$$\delta(^{124}\text{Sn}) - \delta(^{112}\text{Sn}) = 0.194 - 0.107 = 0.087$$

- However, in the Ca isotopes,

$$\delta(^{48}\text{Ca}) - \delta(^{40}\text{Ca}) = 0.167 - 0 = 0.167$$

- Thus, studies of $^{40-48}\text{Ca}$ might provide a more precise determination of K_{sym}



- Comparison of ^{24}Mg excited states obtained in this work with the ones reported in NNDC and by Kawabata *et al.* 2010. The last column shows the difference between calibrated excited states and assigned excited states.

J^π	NNDC (MeV)	Kawabata (MeV)	This work (MeV)	Difference (MeV)
2^+	9.00334	-	9.004	0.0006
1^-	9.14599	-	9.148	0.002
0^+	9.30539	-	9.304	-0.0014
2^+	10.3605	-	10.364	0.0035
0^+	10.6797	-	10.684	0.0043
2^+	11.5182	-	11.517	-0.0012
0^+	11.7281	-	11.724	-0.0041
1^-	11.8649	-	11.864	-0.0009
0^+	13.8500	-	13.854	0.004
4^+	-	15.2500	15.260	0.010
2^+	-	17.6400	17.634	-0.006



