# "LiquidO" An Opaque Detector

### Jeff Hartnell

For the LiquidO Proto-collaboration

US University of Sussex

Conference on Neutrinos and Nuclear Physics, Cape Town 25<sup>th</sup> February 2020

### Outline

- LiquidO
  - Opaque scintillator
  - Lattice of fibre optic cables
- High resolution imaging capabilities
  - Particle identification (PID)
- First papers
- "NoWaSH" scintillator
- Proof of principle experiment
- Physics with LiquidO

### Transparent scintillator detectors A workhorse of neutrino physics





### Opacity

- Two ways to make something opaque
  - Short scattering length
  - Short absorption length







Stochastic confinement of light In a scintillator with a short scattering length the light stays near where it is produced – Each photon undergoes a random walk



#### So, how do we measure the light?



### Lattice of Wavelength Shifting Fibres



## Collect and extract the light from near the point of production using WLS fibres



### LiquidO: Opaque Scintillator + fibres Confine light locally $\rightarrow$ freeze information



Side-on or top-down view

#### **Potential for high-resolution imaging detector**

(without manual segmentation)

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Simple example detector

1 cm pitch lattice of fibres, along a single direction

x [cm] (a) 50 50 e<sup>+</sup> lonization y [cm] 0 Annihilation  $\gamma$  1 Annihilation  $\gamma$  2 -50 40 z [cm] 50 -50 50 x [cm]

Positron of 1 MeV

Two annihilation gammas

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# Use high-resolution imaging for particle identification



Low energy: 2 MeV

## Distinguish positrons from point-like energy depositions (e.g. electrons, protons, alphas)



# Use high-resolution imaging for particle identification





### Use high-resolution imaging for particle identification



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### **First Papers**

#### Neutrino Physics with an Opaque Detector

A. Cahrens<sup>+1,3,10</sup>, A. Abusiems<sup>13</sup>, J. dos Aujos<sup>13</sup>, T. J. C. Bezerra<sup>18</sup>, M. Bougrand<sup>9</sup>, C. Bourgeois<sup>9</sup>, D. Breten<sup>9</sup>, C. Buck<sup>12</sup> J. Basto<sup>4</sup>, E. Calou<sup>5</sup>, E. Chauveau<sup>4</sup>, M. Chen<sup>10</sup>, P. Chimenti<sup>11</sup>, F. Dal Corso<sup>13</sup>, G. De Conto<sup>11</sup>, S. Dusini<sup>1</sup> G. Fiorentini<sup>(5</sup>a<sup>7b</sup>), C. Frigerio Martins<sup>11</sup>, A. Givandan<sup>1</sup>, P. Govoni<sup>10,2b</sup>, B. Gramlich<sup>14</sup>, M. Grassi<sup>1,9</sup>, Y. Han<sup>1,9</sup> J. Hartnell<sup>19</sup>, C. Hugon<sup>6</sup>, S. Jiménez<sup>5</sup>, H. de Kerret<sup>13</sup>, A. Le Newl<sup>8</sup>, P. Loaiza<sup>9</sup>, J. Maalmi<sup>9</sup>, F. Mantowan<sup>76</sup> L. Manzanillas<sup>9</sup>, C. Marquet<sup>4</sup>, J. Martino<sup>18</sup>, D. Navas<sup>5</sup>, H. Nanokawa<sup>14</sup>, M. Obolensky<sup>1</sup>, J. P. Othoa-Ricoux<sup>6,15</sup> G. Ortona<sup>20</sup>, C. Palomares<sup>5</sup>, F. Pessina<sup>11</sup>, A. Pin<sup>4</sup>, M. S. Pravikoff<sup>4</sup>, M. Roche<sup>4</sup>, B. Roskovse<sup>4</sup>, N. Roy<sup>5</sup>, C. Santos<sup>1</sup> A. Seralm<sup>(1),10</sup>, L. Simard<sup>9</sup>, M. Sati<sup>24,20</sup>, L. Stanou<sup>11</sup>, V. Strati<sup>24,70</sup>, J.-S. Stutzmann<sup>10</sup>, F. Sorkane<sup>(1),17</sup>, A. Verdugo<sup>1</sup> B. Viand<sup>18</sup>, C. Volpe<sup>1</sup>, C. Vrignon<sup>1</sup>, S. Wagner<sup>1</sup>, and F. Yermia<sup>18</sup> APC, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité University, 75205 Paris Cedex 13, France <sup>2n</sup>Università di Milano-Bicorca, 1-20126 Milano, Italy <sup>th</sup>INFN, Sezione di Milano-Bicorca, I-20126 Milano, Italy <sup>3</sup>Centro Brasileiro de Pesquinas Físicas (CBPF), Rio de Janeiro, RJ, 22290-180, Brazil <sup>6</sup>CENBG, UMR5797, Université de Bordeaux, CNRS/IN2P3, F-33170, Gradignan, France CIEMAT, Centro de Investigaciones Energéticas, Medicambientales y Tecnológicas (CIEMAT), E-28040 Madrid, Spain <sup>8</sup>Ais Manseille Univ, CNRS/IN2PJ, CPPM, Manseille, France <sup>To</sup>Department of Physics and Earth Sciences, University of Ferrara, Via Saragat 1, 44122 Ferrara, Italy 7hINFN, Ferrara Section, Via Saragat 1, 44122 Ferrara, Italy \*Department of Physics and Astronomy, University of California at Irvine, Irvine, California 92697, USA <sup>9</sup>LAL, Univ. Paris-Sud. CNRS/IN2P3, Université Paris-Saclay, Orsay, France 19LNCA Underground Laboratory, CNRS/IN2P3 - CEA, Chooz, France <sup>11</sup>Departamento de Fúsica, Universidade Estadual de Londrina, 86051-990, Londrina – PR, Brazil <sup>12</sup>Max-Planck-Institut für Kernphysik, 60117 Heidelberg, Germany <sup>13</sup>INFN, Sezione di Padova, via Marzolo 8, I-35131 Padova, Italy <sup>14</sup>Department of Physics, Pontificia Universidade Católica de Rio de Janeiro, Rio de Janeiro, RJ, 22451-900, Ileazii <sup>15</sup>Pontificia Universidad Católica de Chile, Santiago, Chile <sup>10</sup>Department of Physics, Engineering Physics & Astronomy, Queen's University, Kingston, Ontario K7L3N6, Canada. <sup>17</sup> RCNS, Tuhoku University, 6-3 AzzAoba, Aramaki, Aoba-ku, 980-8578, Sendai, Japan <sup>10</sup>SUBATECH, CNRS/IN2P3, Université de Nantes, IMT-Atlantique, 44307 Nantes, France <sup>19</sup>Department of Physics and Astronomy, University of Sussex, Falmer, Brighton BN1 9QR, United Kingdom <sup>20</sup>INFN, Sexiate di Torino, I-10125 Torino, Italy

August 9, 2019

iv:1908.02859v1 [physics.ins-det] in 1956 revolutionised our understanding of the uni- confining and collecting light near its creation point verse at its most fundamental level and provided a with an opaque scintillator and a dense array of finew probe with which to explore the cosmos. Fur- bres. The principles behind LiquidO's detection techthermore, it laid the groundwork for one of the most nique and the results of the first experimental validasuccessful and widely used neutrino detection tech- tion are presented. The LiquidO technique provides nologies to date: the liquid scintillator detector. In high-resolution imaging that enables highly efficient these detectors, the light produced by particle inter- identification of individual particles event-by-event. actions propagates across transparent scintillator vol- Additionally, the exploitation of an opaque medium umes to surrounding photo-sensors. This article in- gives LiquidO natural affinity for using dopants at untroduces a new approach, called LiquidO, that breaks

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Blater Parihal Chaire Fellow

7 Aug 20

The discovery of the neutrino by Reines & Cowan with the conventional paradigm of transparency by precedented levels. With these and other capabilities, LiquidO has the potential to unlock new opportunities in neutrino physics, some of which are discussed here.

Rectments: Assess 72, 2079 Bayessa: September 33, 2079 Accarnia: Deadure 21, 2019 Penannan November 5, 2079

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#### Novel opaque scintillator for neutrino detection

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ABATRACT: There is rising interest in organic scintillators with low scattering length for future neutrino detectors. Therefore, a new scintillator system was developed based on admixtures of paraffin wax in linear alkyl hensene. The transparency and viscosity of this gel-like material can be tuned by temperature adjustment. Whereas it is a colorless transparent liquid at temperatures around 40°C, it has a milky was structure below 20°C. The production and properties of such a scintillator as well as its advantages compared to transparent liquids are described.

Keywowns: Detector design and construction technologies and materials; Neutrino detectors; Scintillators, scintillation and light emission processes (solid, gas and liquid scintillators)

AnXiv aParet: 1908.03334

Contropositing authors

#### More information about LiquidO can be found in arXiv:1908:02859 and arXiv:1908.03334

(See also seminar at CERN: https://indico.cern.ch/event/823865/)

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### What is the Scintillator Mixture?



#### Linear Alkyl Benzene + PPO

Paraffin

Opacity depends on paraffin concentration, which changes crystalisation temperature



"NoWaSH"

### "NoWaSH" Opaque Scintillator



#### Behaviour similar to candle wax Transparent liquid when warm (>36 degC) Opaque solid when cooled (~25 degC)



### Wax (random) structure







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### **Proof of Principle**



#### Three scintillators compared

The predicted and yet remarkable feature:

More light was collected by the WLS fibres near the light source when the scintillator became opaque

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### Next step... "mini-LiquidO"

### Data imminent





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### What physics can we do?

- Lots of ideas
  - Not enough time to cover them all
- Multiple papers in preparation
  - Ultra precise  $\theta_{13}$  measurement for unitarity test
    - https://indico.cern.ch/event/577856/contributions/3421609/attachments/1878723/3095931/190711-UnitariryLiquidOChooz\_EPS-Anatael.pdf
  - <sup>40</sup>K geo-neutrino measurement
    - <u>https://indico.cern.ch/event/825708/contributions/3550280/attachments/1931708/3199582/Serafini\_NGS19.pdf</u>
  - CPv search using pion DAR beam
    - <u>https://arxiv.org/abs/1807.04731</u>
  - CPv search with GeV-scale pion DIF beam
  - and more...

### What physics can we do?



#### Motivation

The DUNE experiment consists of four 10kton liquid-argon detector modules in separate cryostats. The technology choices for the first three modules have been described in the recent DUNE TDR. A fourth module will complete the DUNE Far Detector. This module provides opportunity for further development of liquid-argon or alternate technologies in support of the DUNE physics goals.

The DUNE Collaboration invites the broader particle physics community to participate in a workshop to explore opportunities for novel detector technologies for this "module of opportunity". The meeting will cover all major aspects of liquid argon TPC technology, including proposals for improvements in tracking, photon detection, electronics, high voltage, electronics and dataacquisition, considering improvements in detector integration and installation.

New detector concepts that can satisfy and expand the DUNE physics goals are encouraged.

#### 

Workshop Dates November 12–13, 2019

Event ID 41055

Workshop Venue Brookhaven National Laboratory Upton, NY 11973 USA

Getting to Brookhaven Lab



### DUNE and 4th Module of Opportunity

• Four separate 17 kt (> 10 kt fiducial) LAr TPCs

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 4 identically sized cryostats: 2 single phase (SP) + 1 dual phase (DP) +1 "opportunity" (this 2+1+1 plan is described in TDR)



### Scalability

- No showstoppers foreseen when scaling LiquidO to ~10 ktons:
  - Invaluable experience from NOvA
  - <u>Key difference</u>: avoid light losses due to reflection inside the cells







A NOvA-sized LiquidO would achieve at least 100 PEs/MeV with today's technology→ **already excellent for MeV physics** 

- Rough cost expected to be comparable to NOvA FD
- Other advantages compared to other detectors:
  - Room temperature operation (no need for cryostat)
  - Self-shielding detector

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### Beautiful LiquidO beam events



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![](_page_24_Figure_0.jpeg)

#### - LiquidO would reveal GeV-neutrino interactions in extremely powerful way:

Imaging capabilities comparable to those of LArTPC Complementary features unique to LiquidO

![](_page_24_Picture_4.jpeg)

### **Additional Physics**

Supernova neutrinos:

- Low energy threshold (~0.1 MeV)
- Channels not accessible with other detectors
- Charge sign ID (e+/e-)
- Directionality information for events ≥ 10 MeV
- Sensitivity to Diffuse Supernova Neutrino Background

#### Search for nucleon decay:

- Very high-efficiency
- Full topological information and sign-ID for some channels through final Michel electron
  - Many exclusive final states

Others (geoneutrinos, reactor antineutrinos... etc)

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![](_page_25_Figure_12.jpeg)

### Doping

- Doped scintillators used with great success
  - Reactor  $\theta_{13}$  experiments (Gadolinium)
  - ZamLAND-Zen (Xe)
- Often a challenge to go to higher loading
  - Can end up making the scintillator opaque

#### Great! ©

- For LiquidO, needing opacity removes the transparency requirement
  - New landscape of novel materials/mixtures to consider
  - It's a whole new (counter-intuitive!) way of thinking about scintillator detectors

### Conclusions

- Innovative new "LiquidO" detector technology
  - Opaque scintillator + lattice of fibres
  - Builds on established technology
- Brings new physics opportunities
  - High-resolution imaging and particle identification
  - Low energy threshold
  - Potential for high levels of doping
- R&D progressing rapidly and steadily

   Proof of principle done. Larger detector data imminent
- LiquidO is a whole new way of thinking about a scintillator detector
  - we've only just scratched the surface

### The End

![](_page_28_Picture_1.jpeg)

### Backup slides

![](_page_29_Picture_1.jpeg)

### Advantages of LiquidO @ DUNE

- Complementary detector properties and capabilities:
  - Low-Z (radiation length 0.5m vs. 0.14m in LArTPC)
  - Self-segmenting detector (no dead material & lower cost)
  - Largest density of free-protons (without being explosive)
  - Low energy threshold
  - Sensitivity to neutrons (scattering and capture)
  - Charge sign ID from Michel e+/e- (separate time scale)
  - High duty cycle and fast timing
- Other opportunities:
  - Plenty of room for optimization depending on physics goals vs. cost

Example of event with 1 cm fibre pitch —

 Room for enhancements such as loading (e.g. Indium) and magnetization

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![](_page_30_Figure_13.jpeg)

![](_page_30_Figure_14.jpeg)

#### Measurement

NoWash 10% & LAB / Fiber 0 (bottom)

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![](_page_31_Figure_2.jpeg)